

# Low-mass WIMP detection with TREX-DM

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*on behalf of TREX-DM team*

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**Centro de Astropartículas y  
Física de Altas Energías**

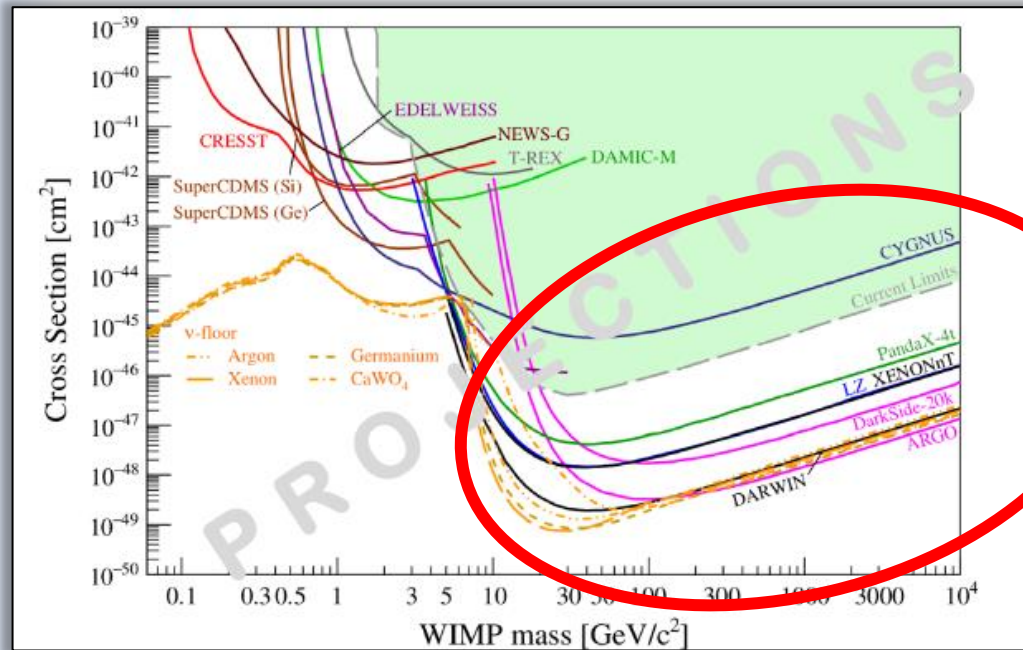
**Universidad Zaragoza**



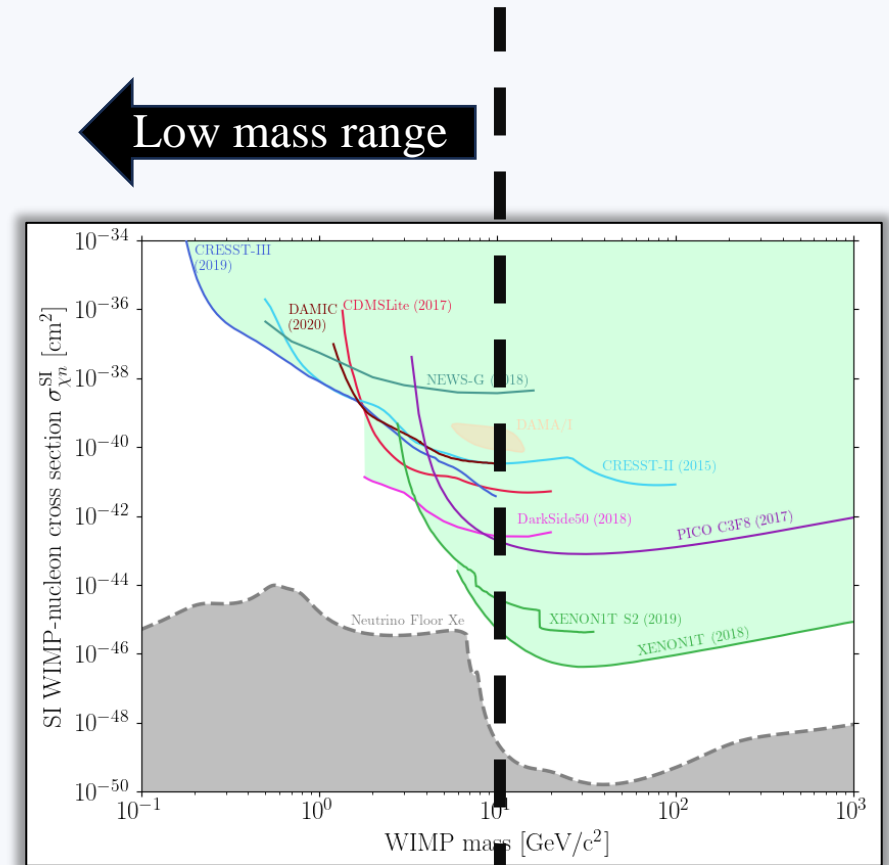
*Laboratorio Subterráneo de Canfranc*

# Goal

- Direct detection of WIMPs (dark matter)
  - Low masses  $< 10 \text{ GeV}/c^2$
- Spin independent nuclear recoil

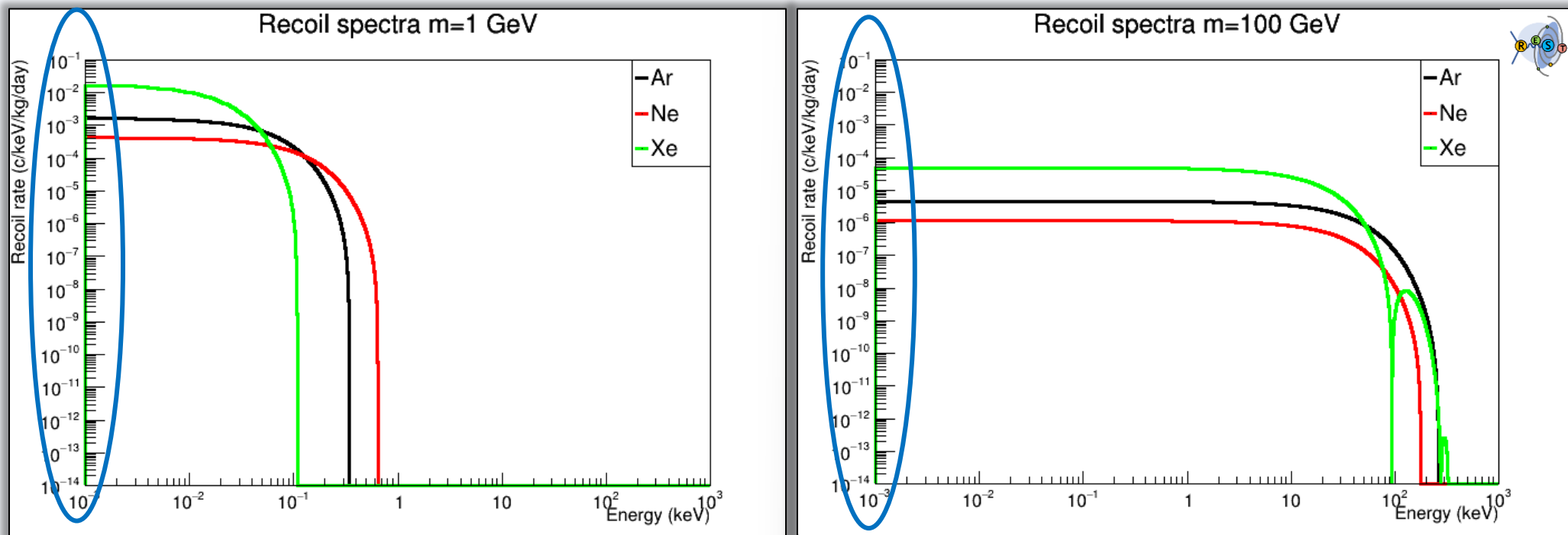


Projected to reach  
neutrino fog in high  
mass soon!



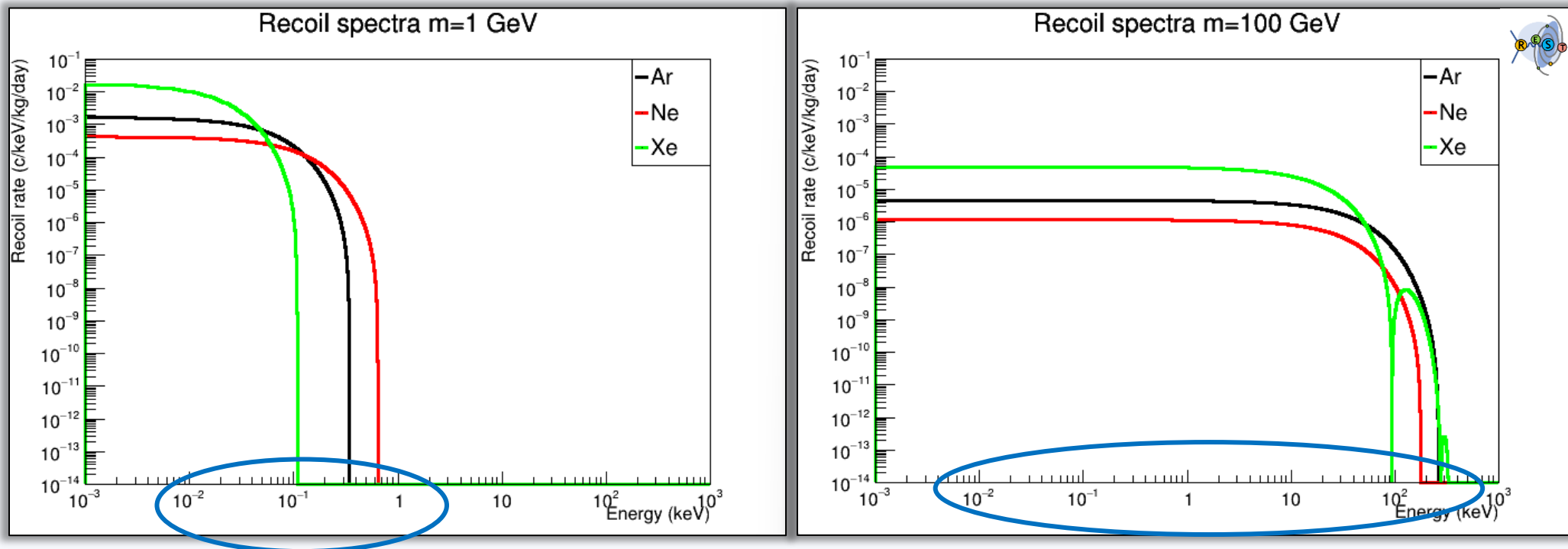
# Challenges

- Low-mass WIMP detection requires:
  - low background
  - high exposure (mass  $\times$  time)
  - low energy threshold
  - light nuclei targets



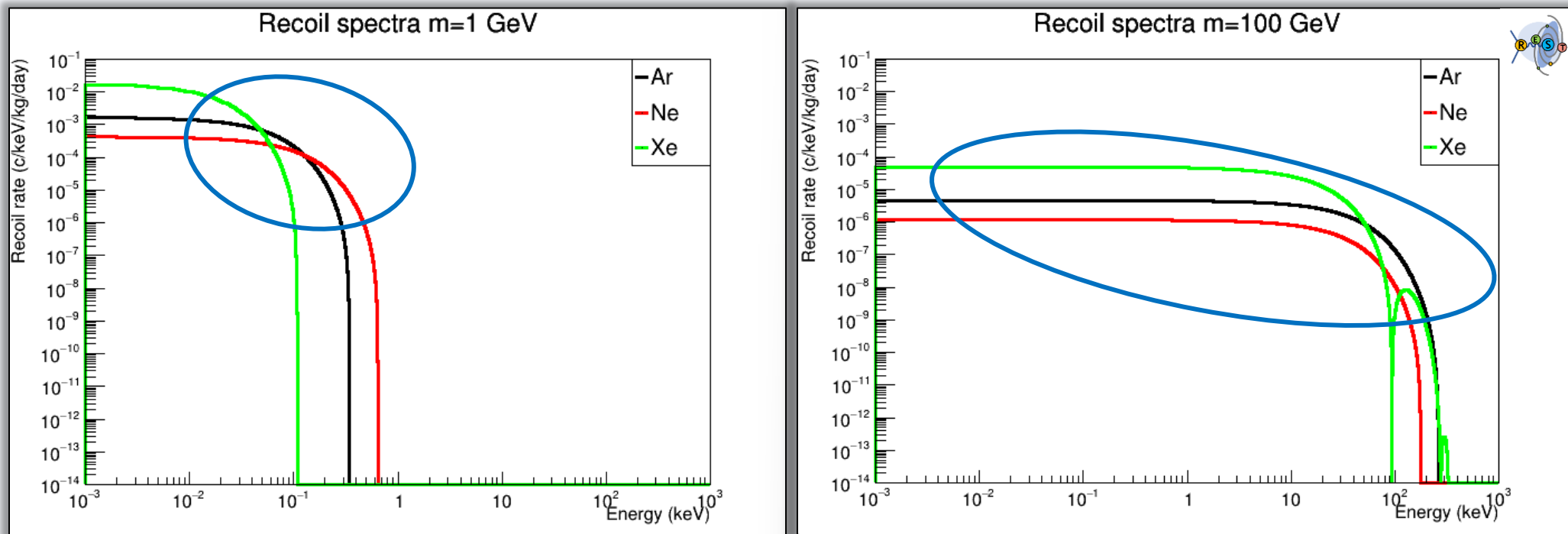
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# Challenges

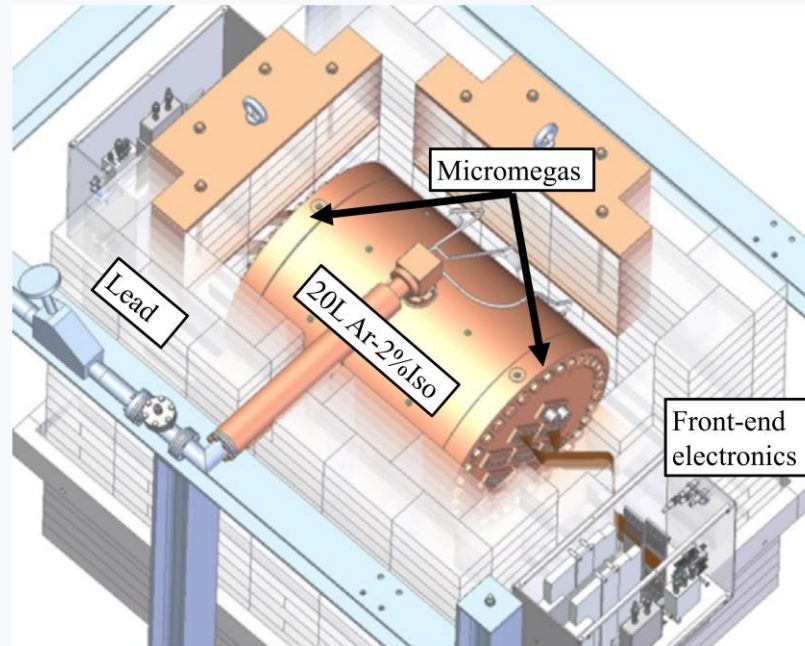
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# TREX-DM specifications

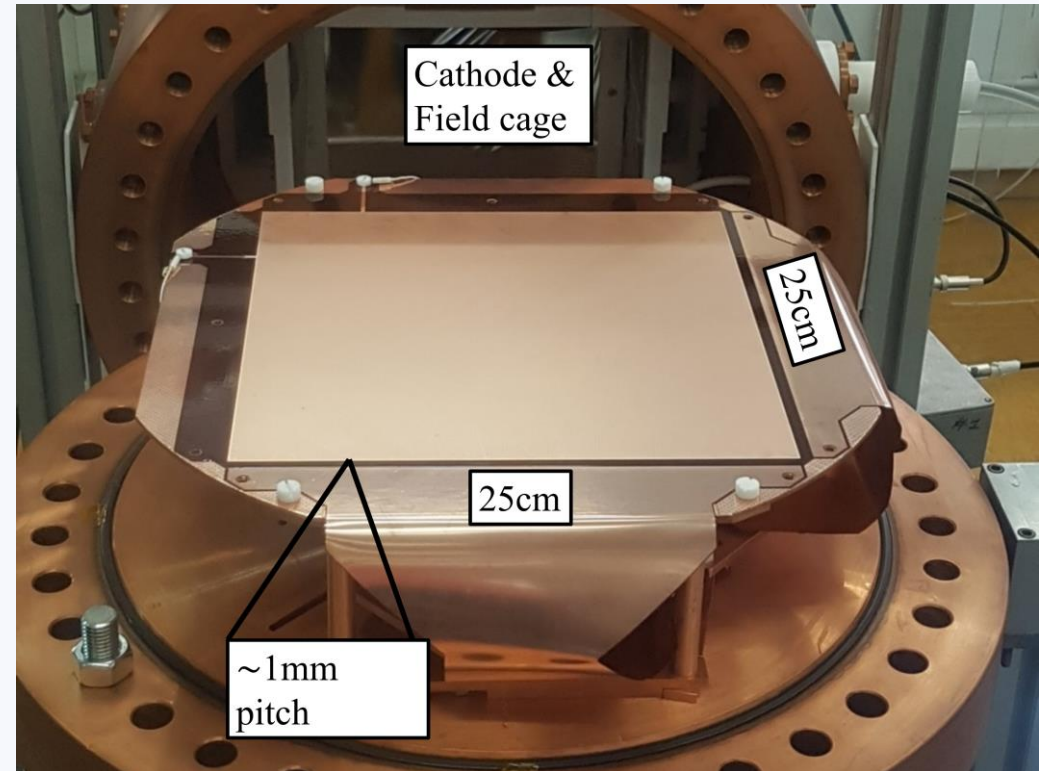
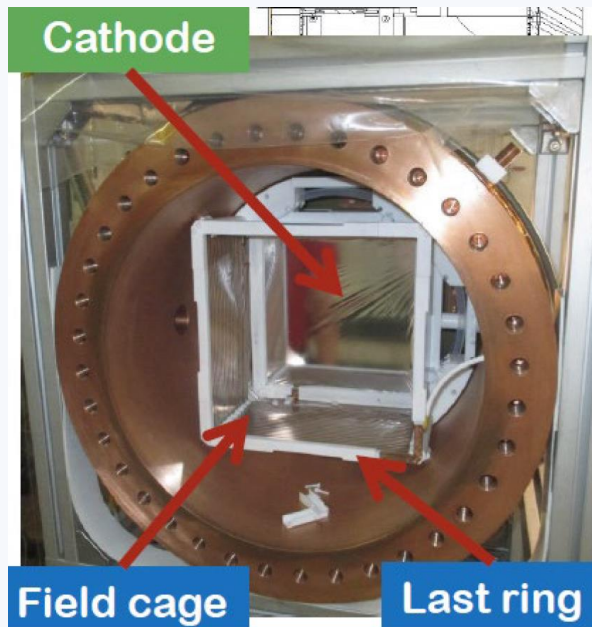
Iguaz, F.J., Garza, J.G., Aznar, F. et al. TREX-DM: a low-background Micromegas-based TPC for low-mass WIMP detection. *Eur. Phys. J. C* 76, 529 (2016). <https://doi.org/10.1140/epjc/s10052-016-4372-6>

- Gaseous TPC of cylindrical copper vessel
- Active volumen of 20L up to 10bar ( **$\sim 0.32$  kg Ar** or  **$\sim 0.16$  kg Ne**)
- Shielding
  - 5cm copper + 20cm lead walls
  - Polyethylene ceiling + water
- Located at LSC (2400 m.w.e.)



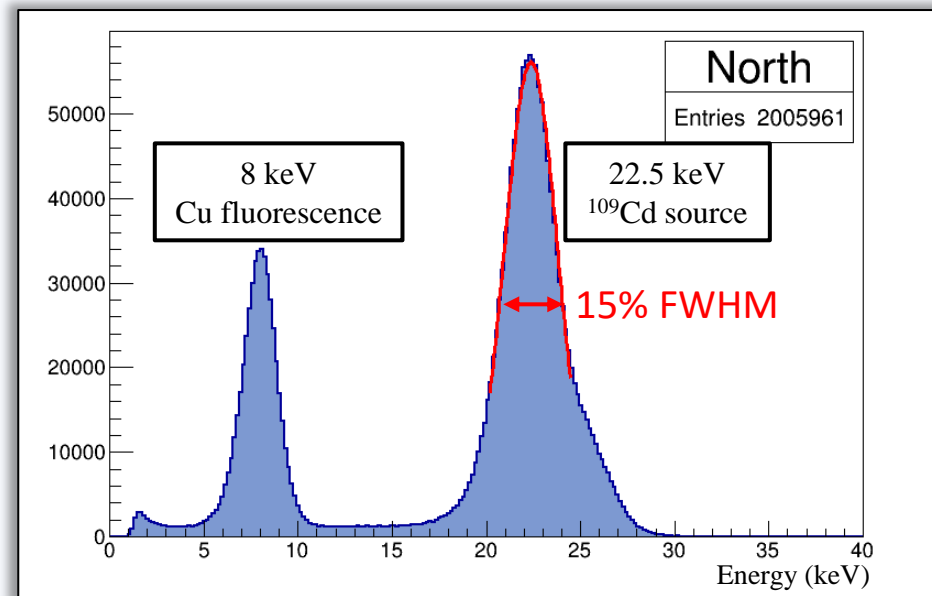
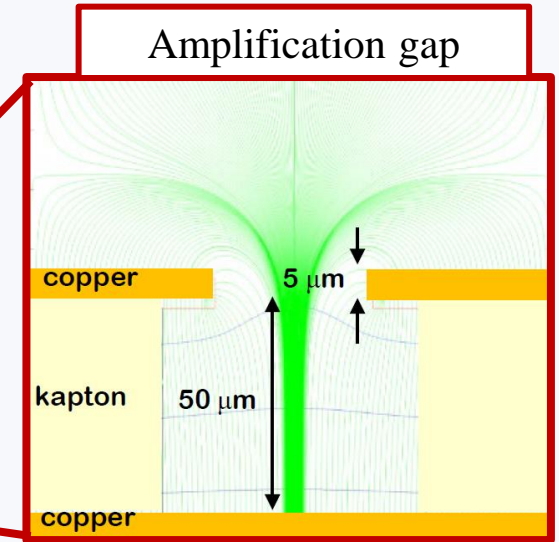
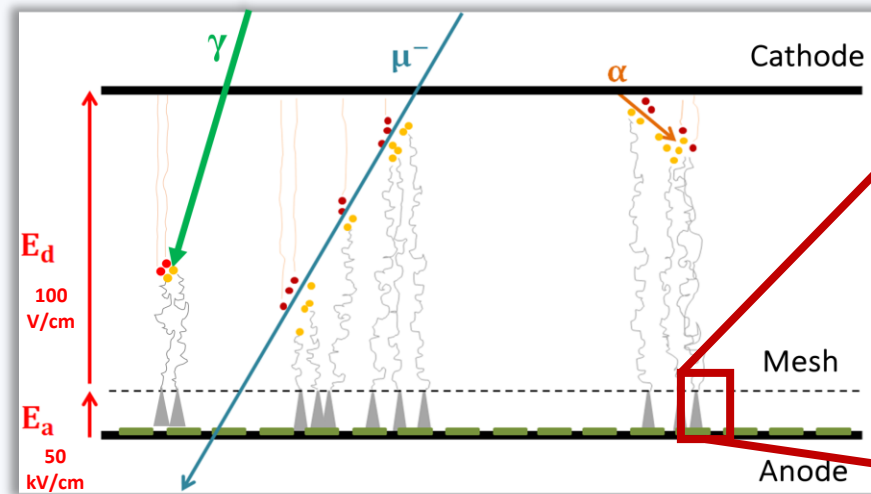
# TREX-DM specifications

- 1 cathode and 2 readout planes
- Largest surface produced (*micro-bulk*) Micro-Mesh Gaseous Structure (**Micromegas**)
  - 512 channels: 256 X strips, 256 Y strips



# Micromegas detector

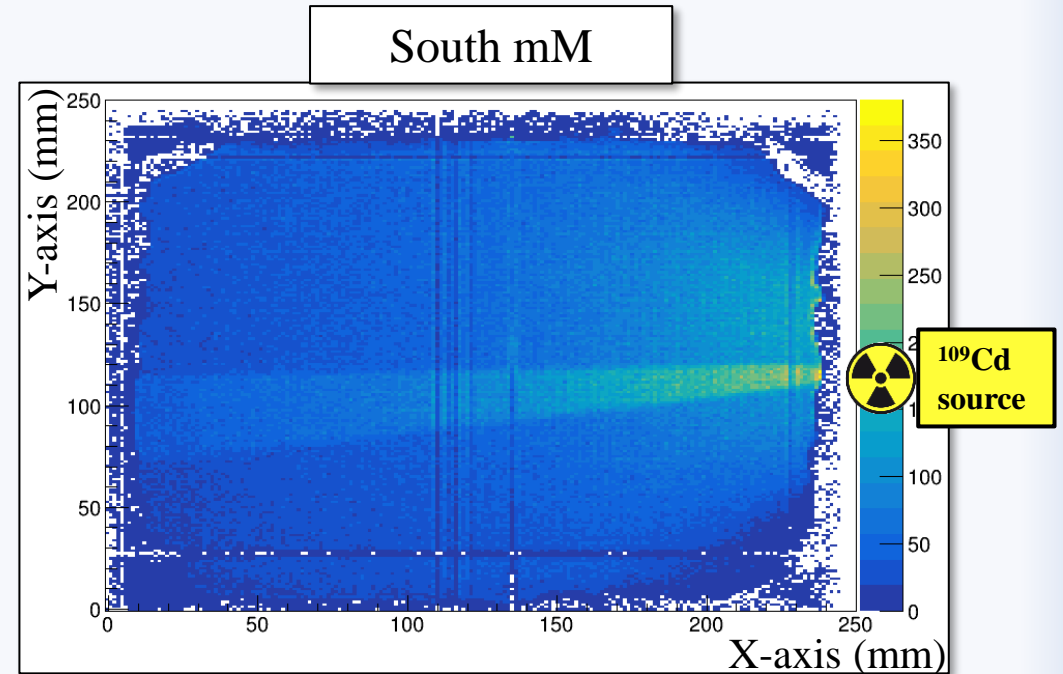
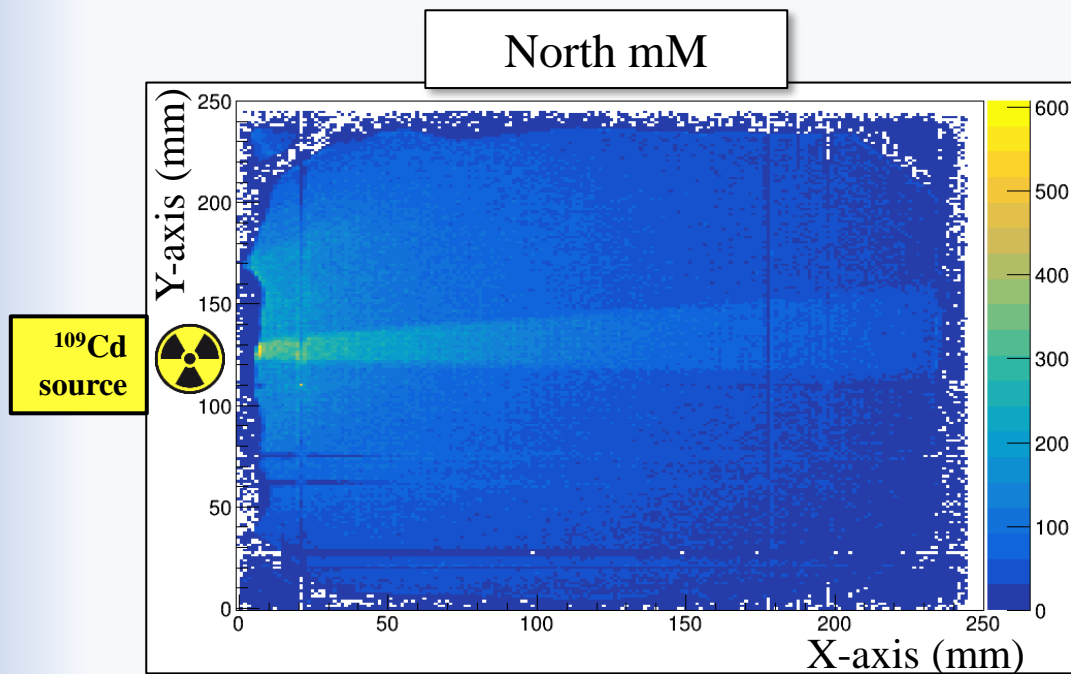
- Working principle of TPC with mM
- Low intrinsic radioactivity
- Absolute gain  $10^3$ - $10^4$





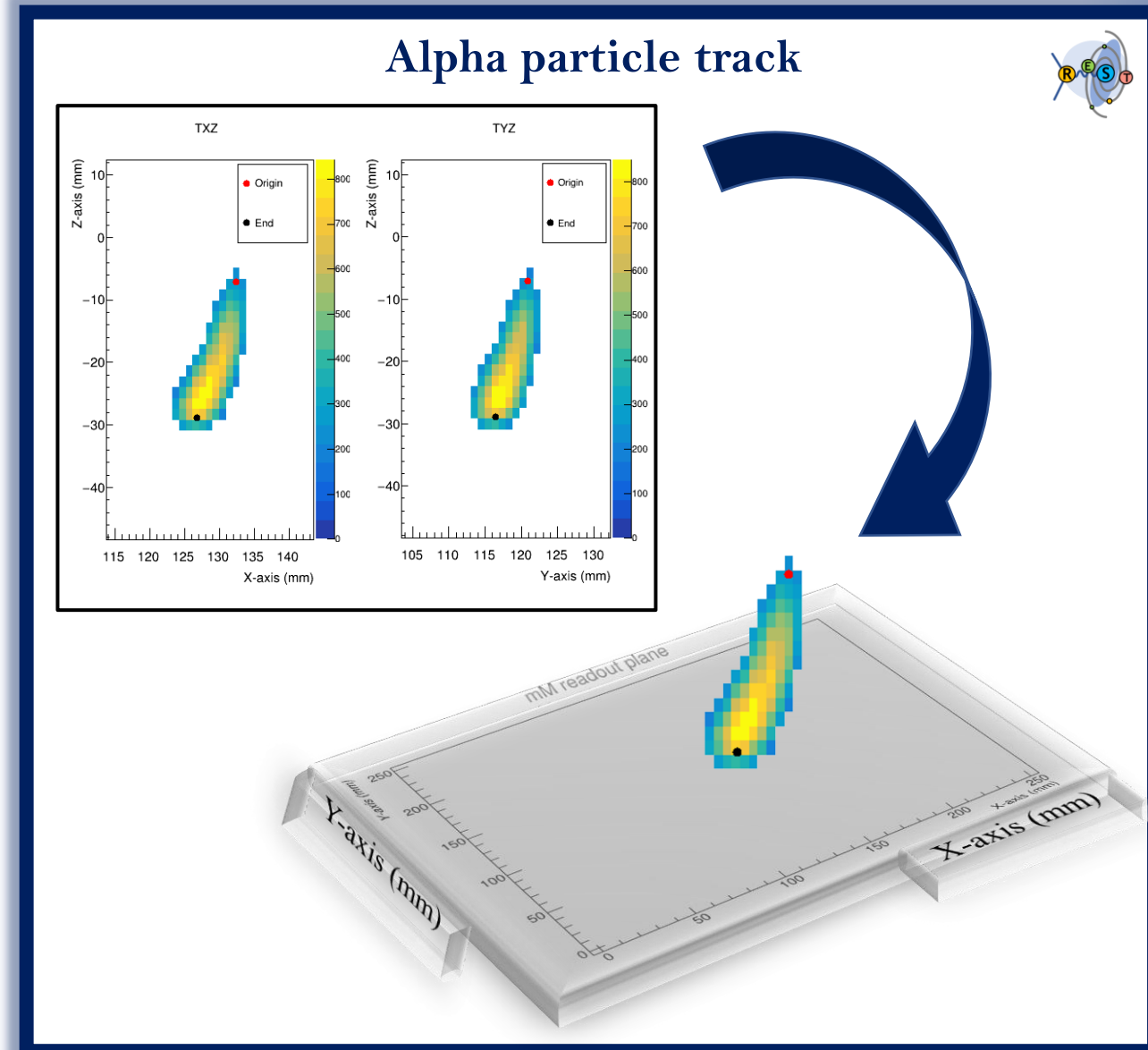
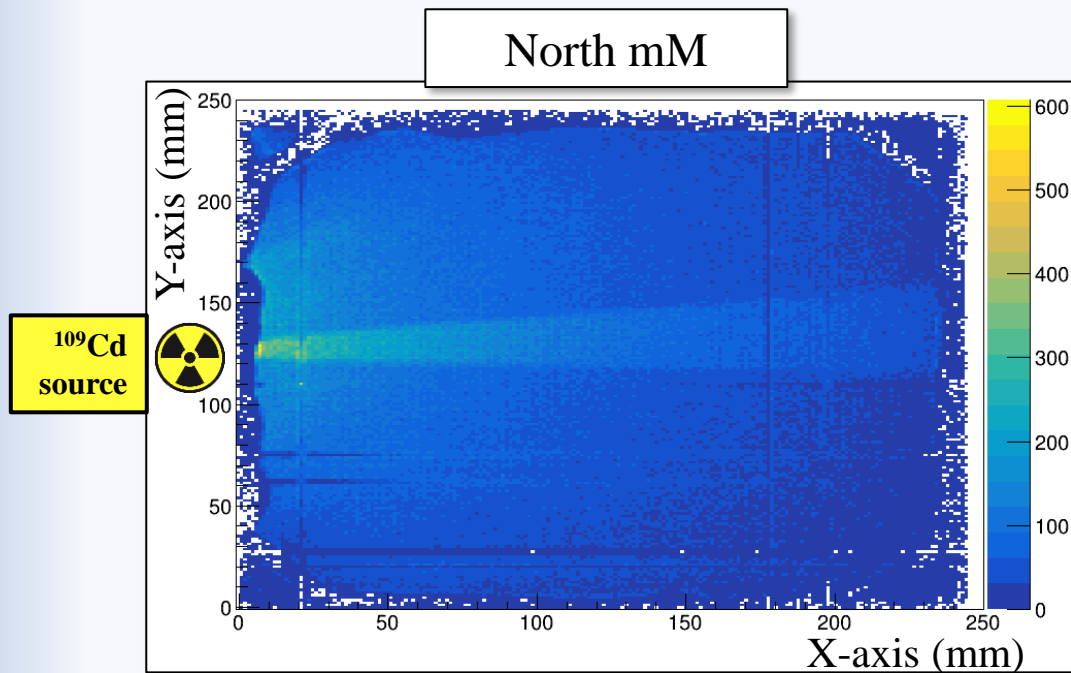
# Micromegas detector

- Topological information
  - background discrimination



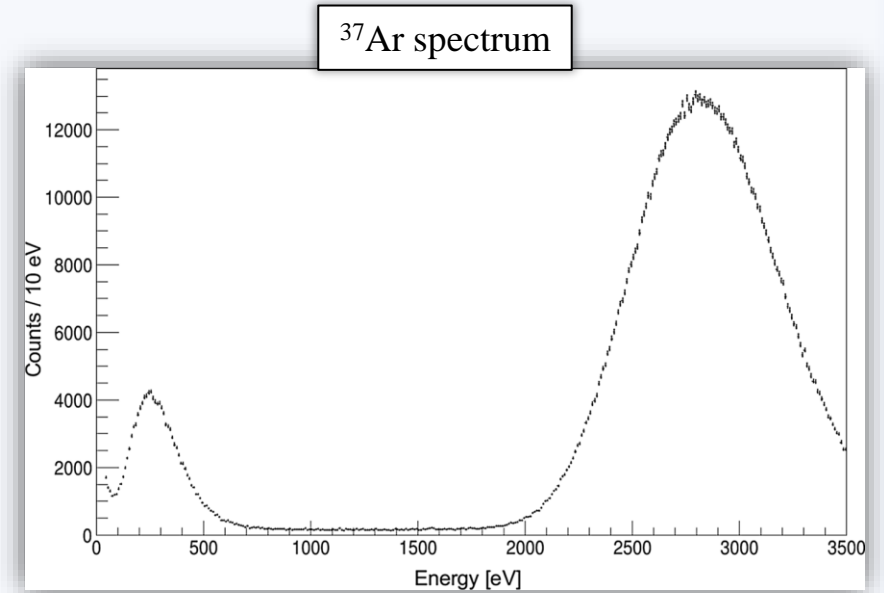
# Micromegas detector

- Topological information
  - background discrimination



# Low energy calibration

- Currently two  $^{109}\text{Cd}$  source are used
- $^{37}\text{Ar}$ 
  - 2.82 keV (K shell, probability 0.90)
  - 0.27 keV (L shell, probability 0.09)
  - Used in XENON1T and NEWS-G
- Gas source  $\rightarrow$  homogeneous illumination
- By  $^{40}\text{Ca}$  irradiation with fast neutrons (AmBe source)

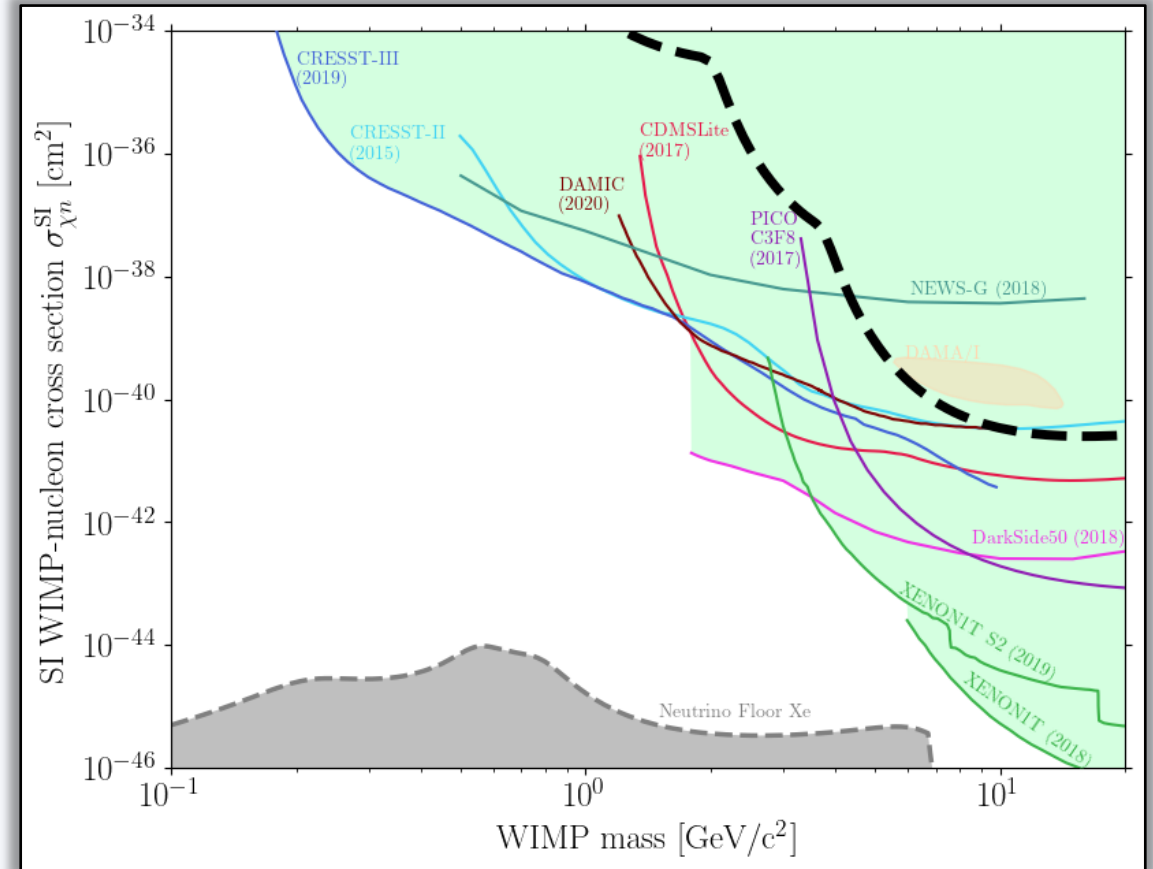


# State of art

	$E_{th}$ (eV <sub>ee</sub> )	B(dru)	Gas
Z	1000	100	Ar-1%Iso

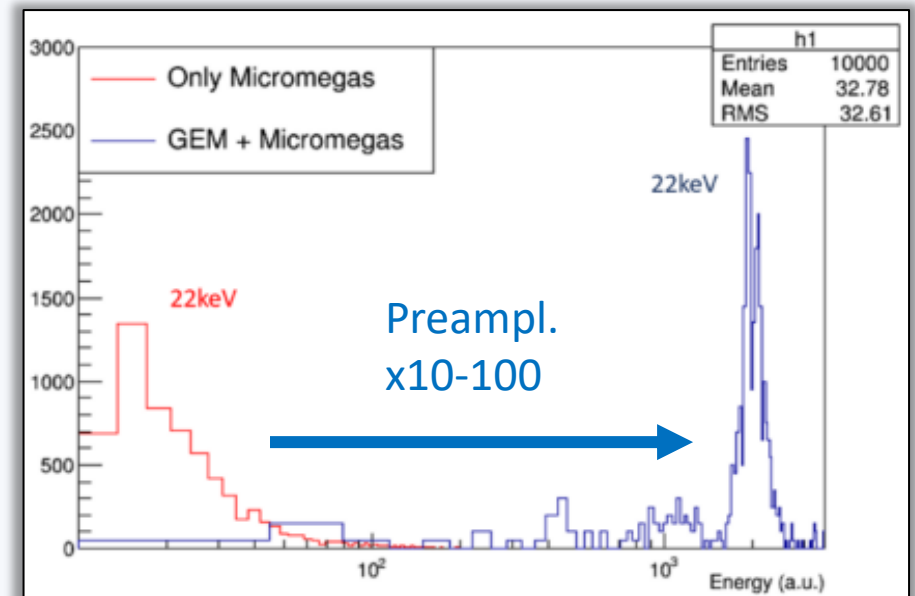
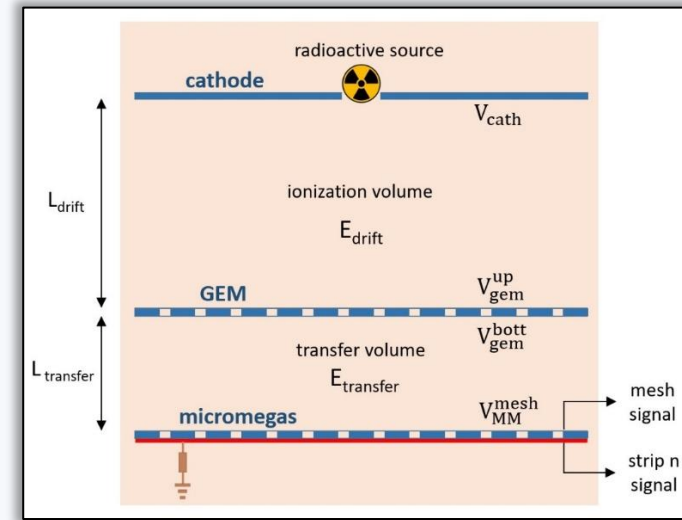
(dru = keV<sup>-1</sup> kg<sup>-1</sup> day<sup>-1</sup>)

Exposure 0.32 kg y



# Improving energy threshold

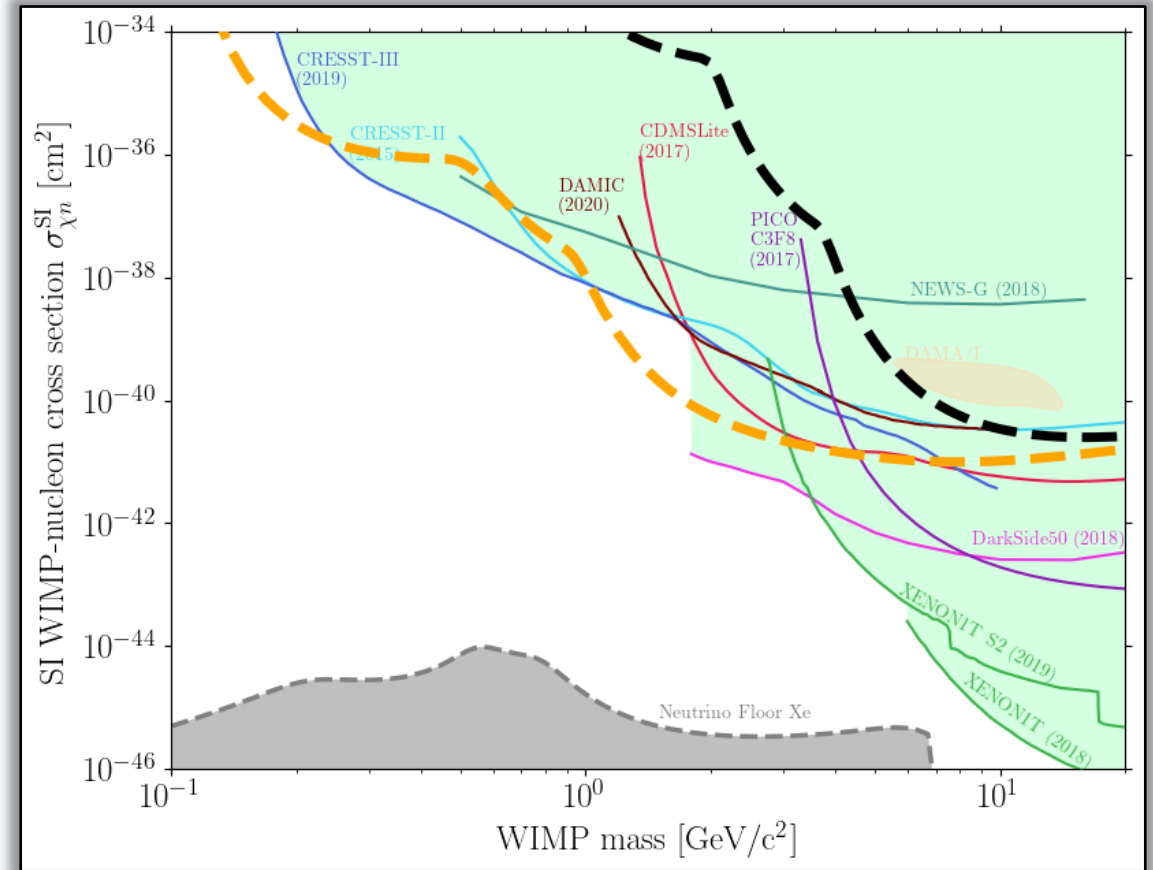
- Improve gain by adding a Gas Electron Multiplier (**GEM**) amplification stage on top of the Micromegas planes
  - increase signal-to-noise ratio
- Preliminary results in test set-up shows x10-100 extra amplification factor
  - $E_{\text{thr}}$  lowered down to single electron ( $20 \text{ eV}_{\text{ee}}$ )



# Sensitivity prospects

	$E_{th}$ (eV <sub>ee</sub> )	B(dru)	Gas
Z	1000	100	Ar-1%Iso
A	50	100	Ar-1%Iso

Exposure 0.32 kg y



# Improving background



## 2019 – Commissioning & Ne data-taking

- I. Ne data-taking with H<sub>2</sub>O & O<sub>2</sub> filters and recirculating
- II. Background dominated by <sup>222</sup>Rn
  - 600 c keV<sup>-1</sup> kg<sup>-1</sup> day<sup>-1</sup> (= dru)
- III. Solution: open-loop operation
  - 100 dru (x10-100 higher than background model)

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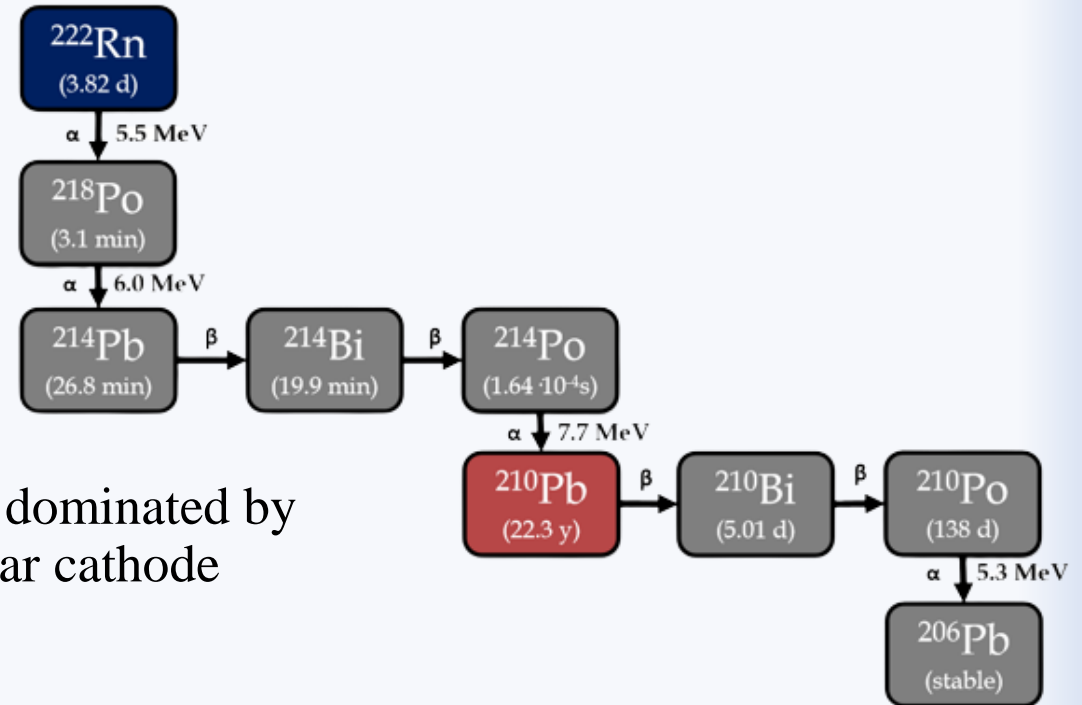
➤ 600 c keV<sup>-1</sup> kg<sup>-1</sup> day<sup>-1</sup> (= dru)

III. Solution: open-loop operation

➤ 100 dru (x10-100 higher than background model)

## 2022 – New & cleaner micromegas installed YouTube

IV. Alpha measurements shows this background is dominated by <sup>222</sup>Rn progeny (<sup>210</sup>Pb) on the surface of the mylar cathode





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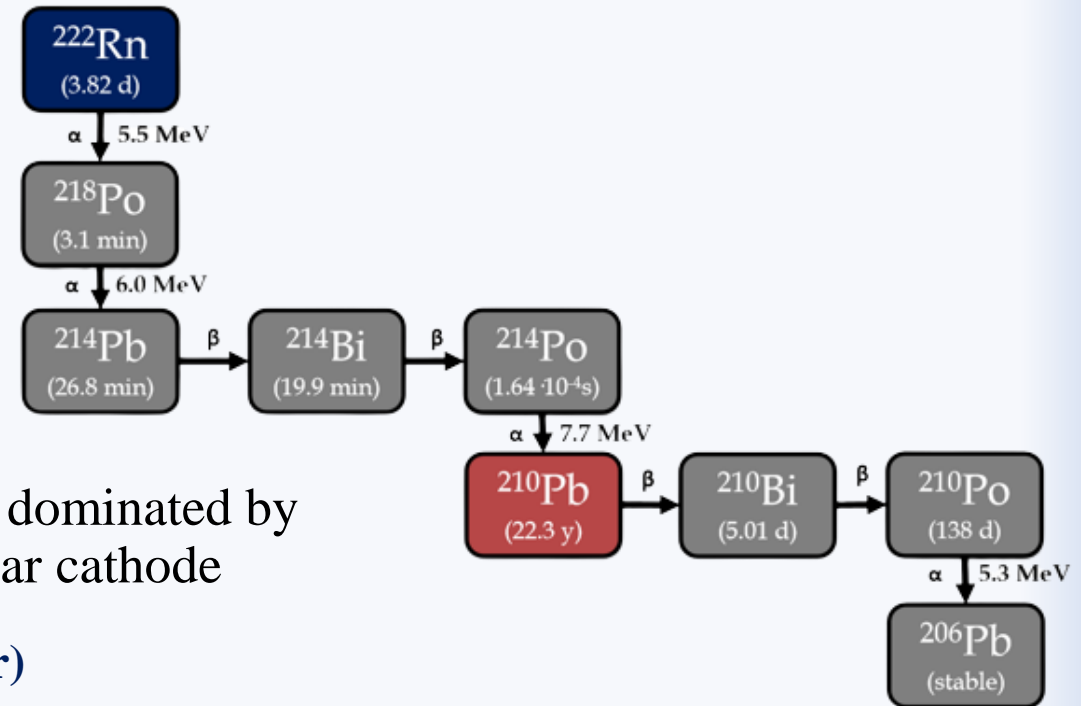
## 2022 – New & cleaner micromegas installed YouTube

IV. Alpha measurements shows this background is dominated by <sup>222</sup>Rn progeny (<sup>210</sup>Pb) on the surface of the mylar cathode

## 2023 – Relocation & Ne not available anymore (change to Ar)

V. Plans to install a more radiopure cathode in the coming months

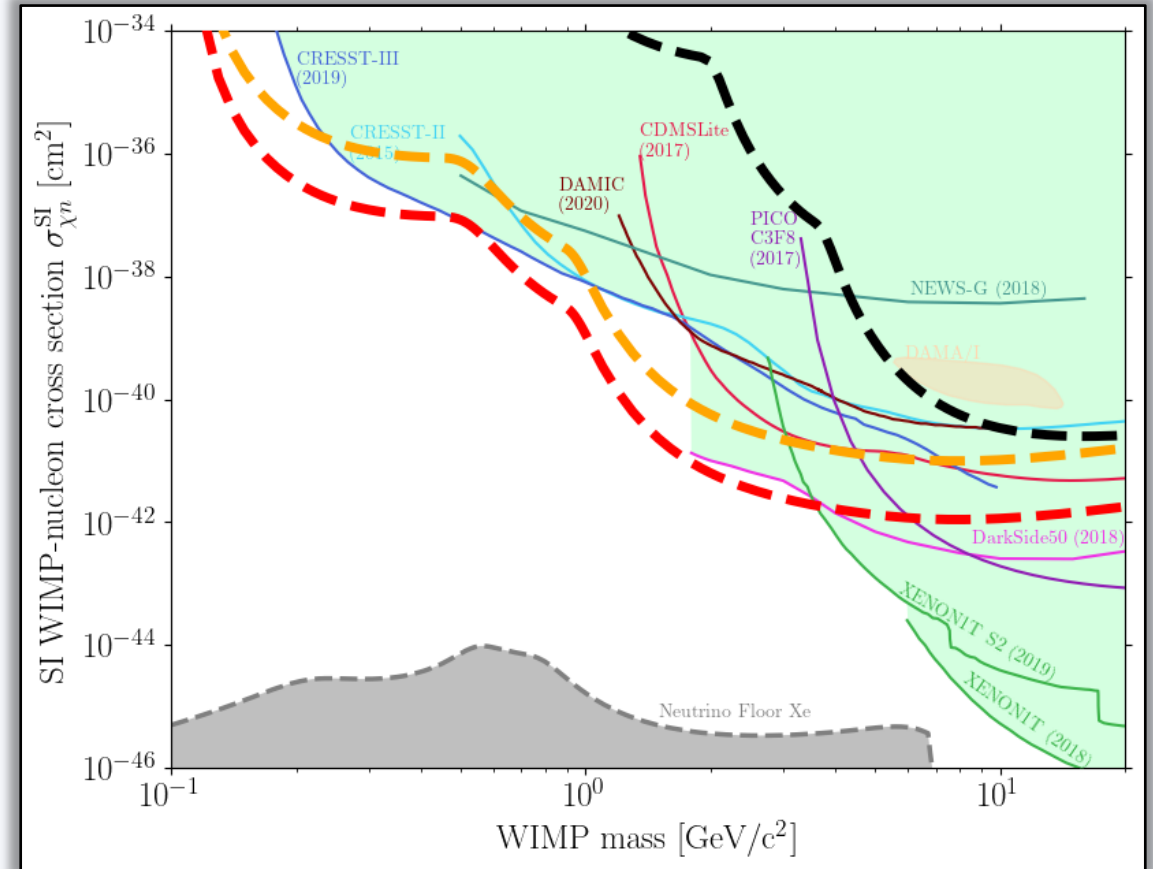
2024 ➤ 1-10 dru



# Sensitivity prospects

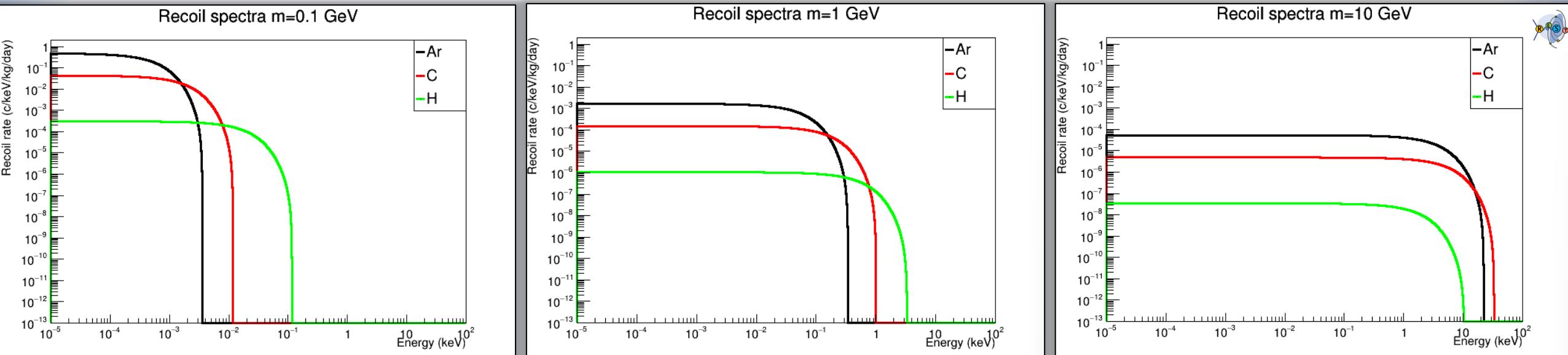
	$E_{th}$ (eV <sub>ee</sub> )	B(dru)	Gas
Z	1000	100	Ar-1%Iso
A	50	100	Ar-1%Iso
B	50	1	Ar-1%Iso

Exposure 0.32 kg y



# Increasing quencher in gas

- For low-mass WIMPs, the lighter the better

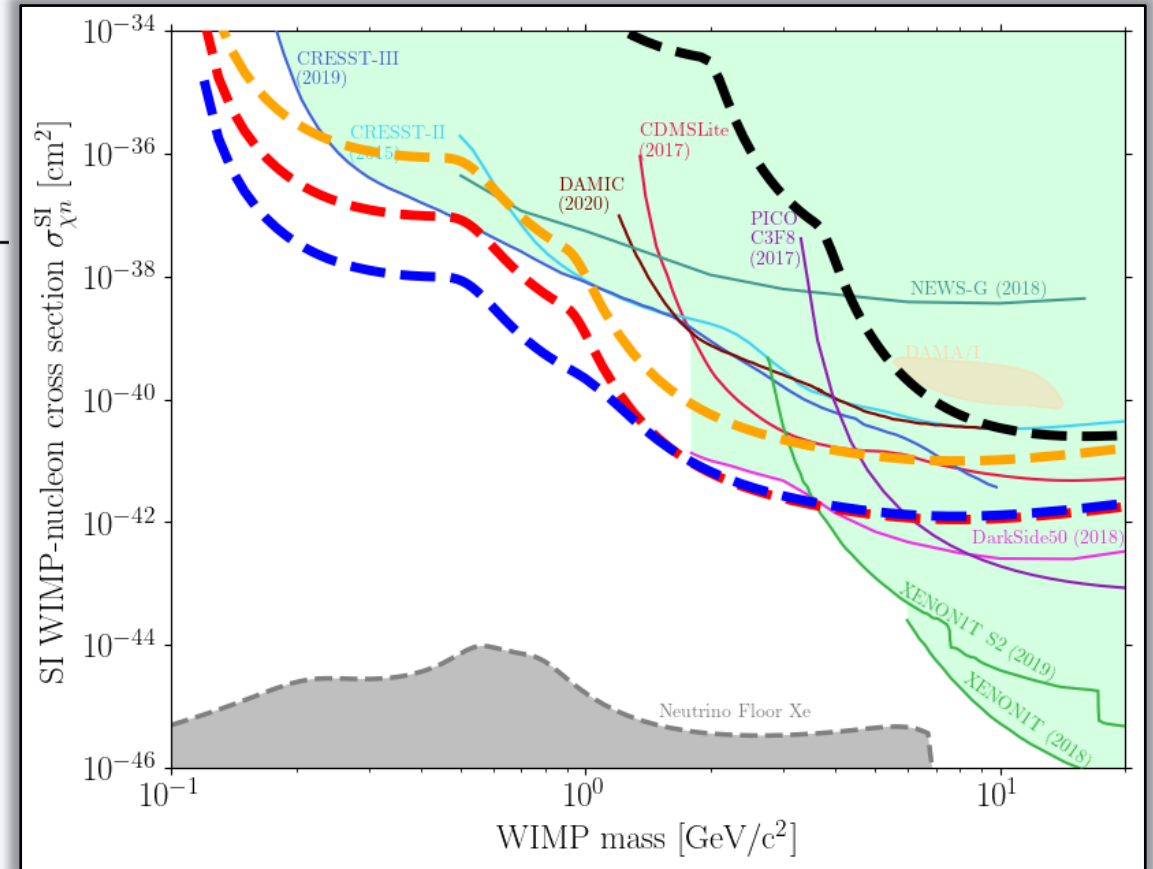


Decreasing WIMP mass

# Sensitivity prospects

	$E_{th}$ (eV <sub>ee</sub> )	B(dru)	Gas
Z	1000	100	Ar-1%Iso
A	50	100	Ar-1%Iso
B	50	1	Ar-1%Iso
C	50	1	Ar-10%Iso

Exposure 0.32 kg y

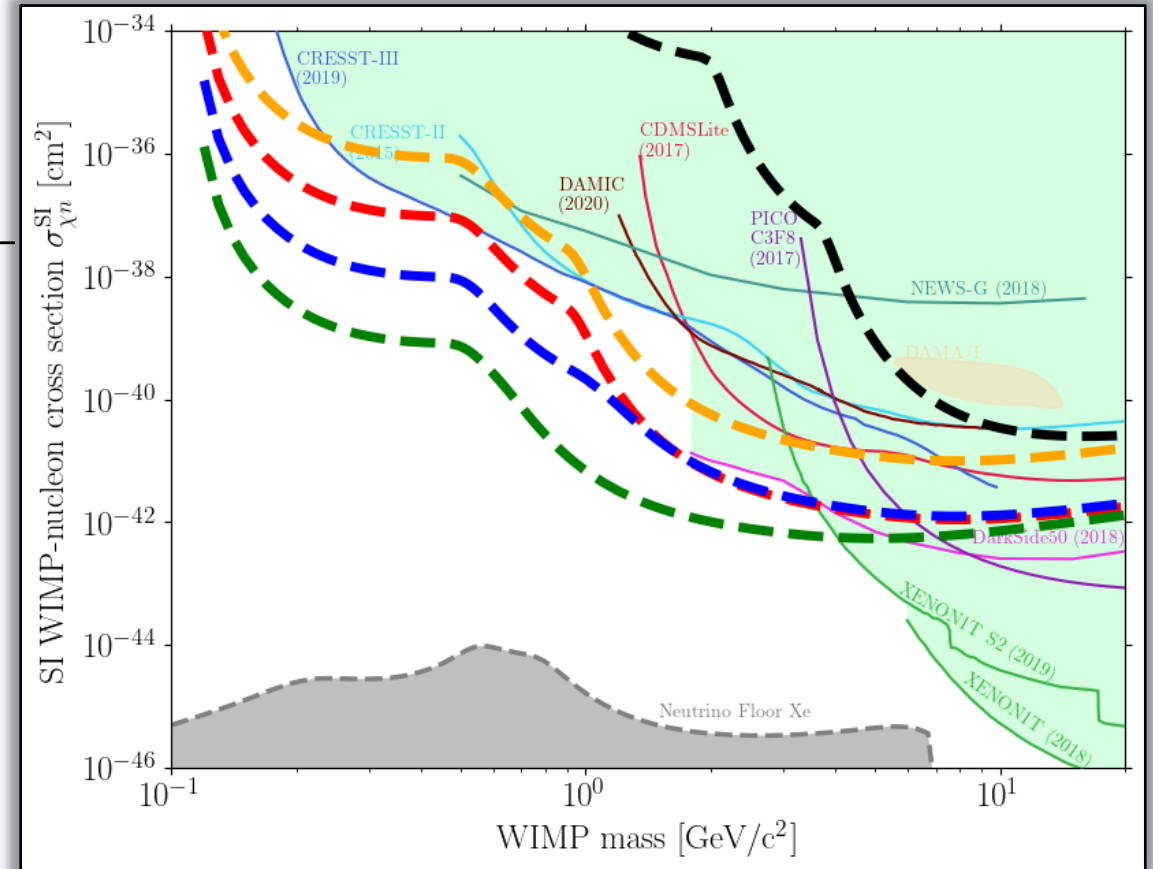


# Sensitivity prospects

	$E_{th}$ (eV <sub>ee</sub> )	B(dru)	Gas
Z	1000	100	Ar-1%Iso
A	50	100	Ar-1%Iso
B	50	1	Ar-1%Iso
C	50	1	Ar-10%Iso
*D	50	0.1	Ne-10%Iso

Exposure 0.32 kg y

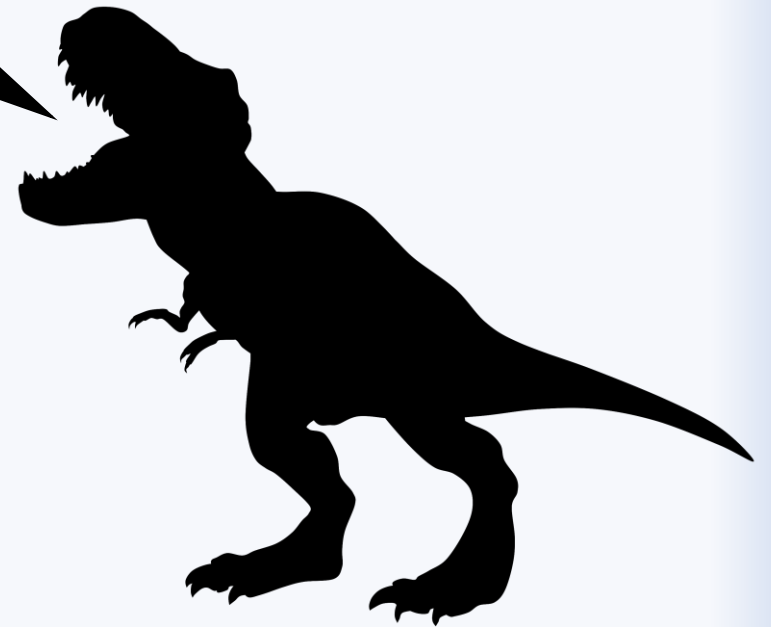
\*Exposure 1.6 kg y



# Summary

- TREX-DM is the only (non-directional) experiment with Micromegas technology searching for (low-mass) WIMPs
  - High radiopurity
  - Topology information → background & noise discrimination
  
- Upgrades planned to be implemented in the near future
  - I. GEM pre-amplification & low-energy calibration
  - II. Cleaner materials (cathode)
  - III. More isobutane

**THANKS!**

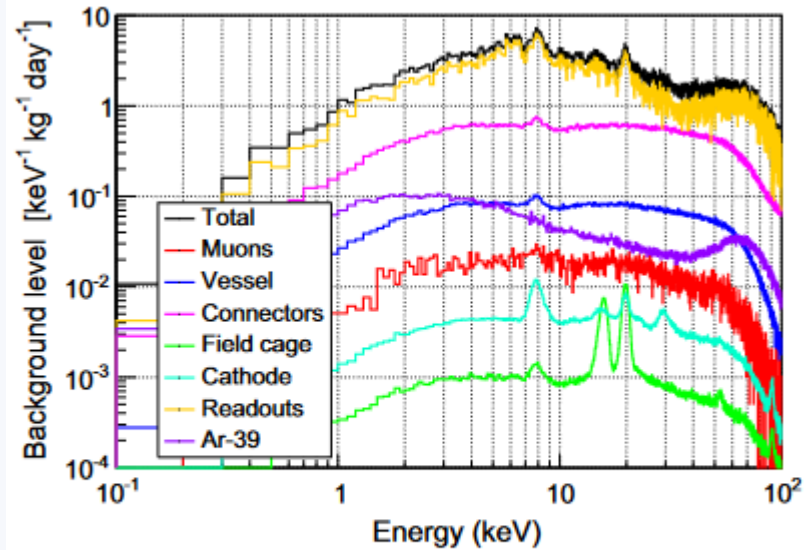


Installation of new  
micromegas

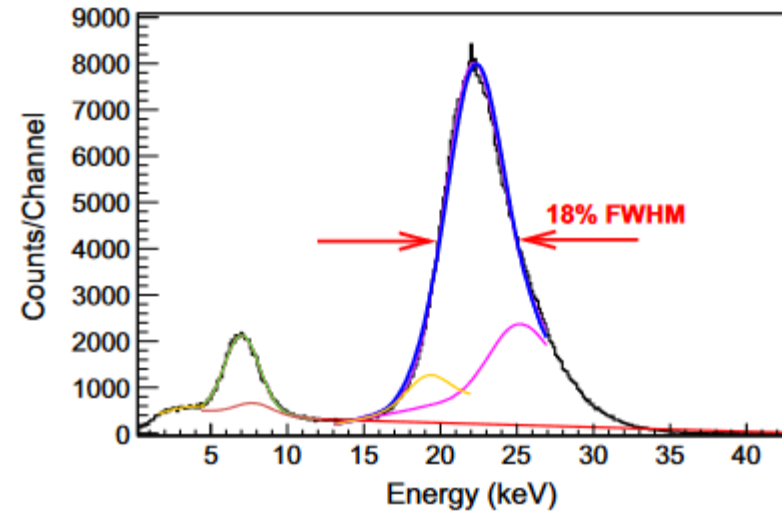


# BACK-UP



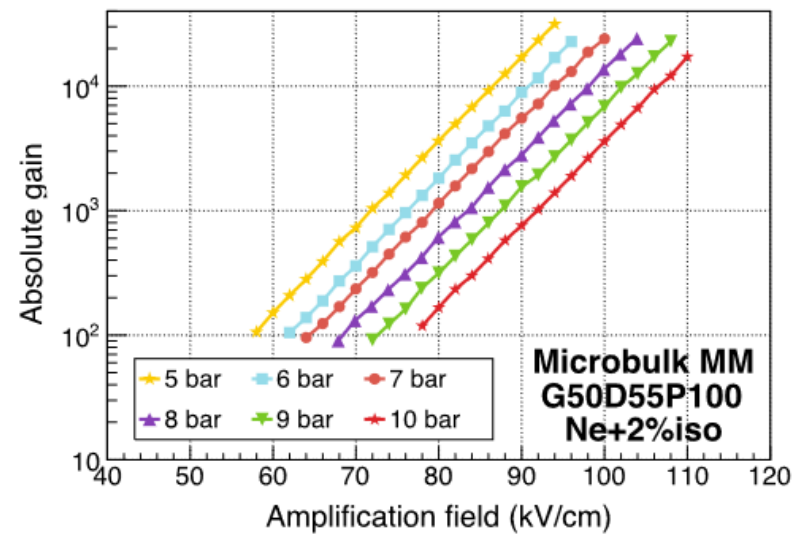
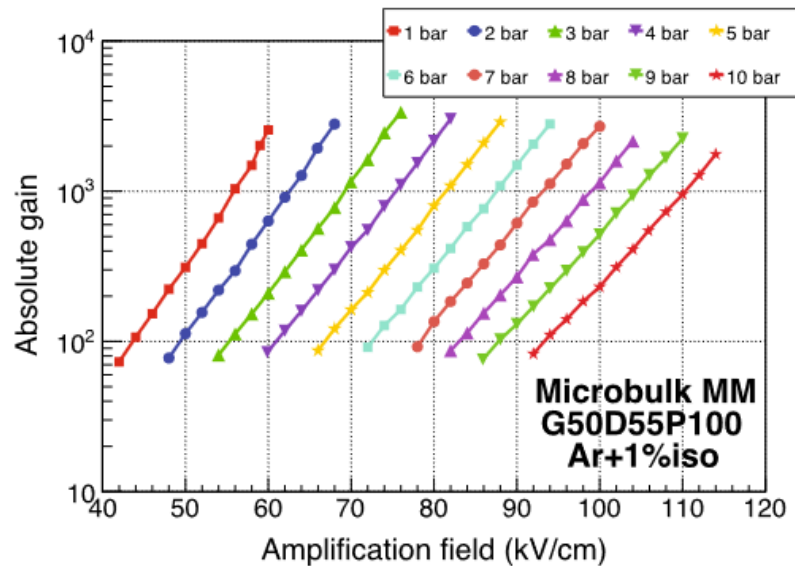


**Fig. 20** Simulated background spectrum expected in TREX-DM experiment (black line) during a physics run at LSC if operated in Ar+2%iC<sub>4</sub>H<sub>10</sub> (left) or Ne+2%iC<sub>4</sub>H<sub>10</sub> at 10 bar. The contribution of



**Fig. 6** Energy spectrum generated by the mesh signals when one of the Micromegas readouts is irradiated by a <sup>109</sup>Cd source in Ar+2%iC<sub>4</sub>H<sub>10</sub> at 2 bar. The spectral parameters are defined through an iterative multi-Gaussian fit corresponding to the K<sub>α</sub> (22.1 keV, blue line) and K<sub>β</sub> (24.9 keV, magenta line) emission lines of the source and their escape peaks (located at 19.1 and 21.9 keV, orange line). The fluorescence lines of iron (at 6.4 keV, emitted from the mesh) and copper (8 keV, from the vessel or the field cage strips) are also present (green and brown lines, respectively)

Iguaz, F.J., Garza, J.G., Aznar, F. *et al.* TREX-DM: a low-background Micromegas-based TPC for low-mass WIMP detection. *Eur. Phys. J. C* **76**, 529 (2016). <https://doi.org/10.1140/epjc/s10052-016-4372-6>

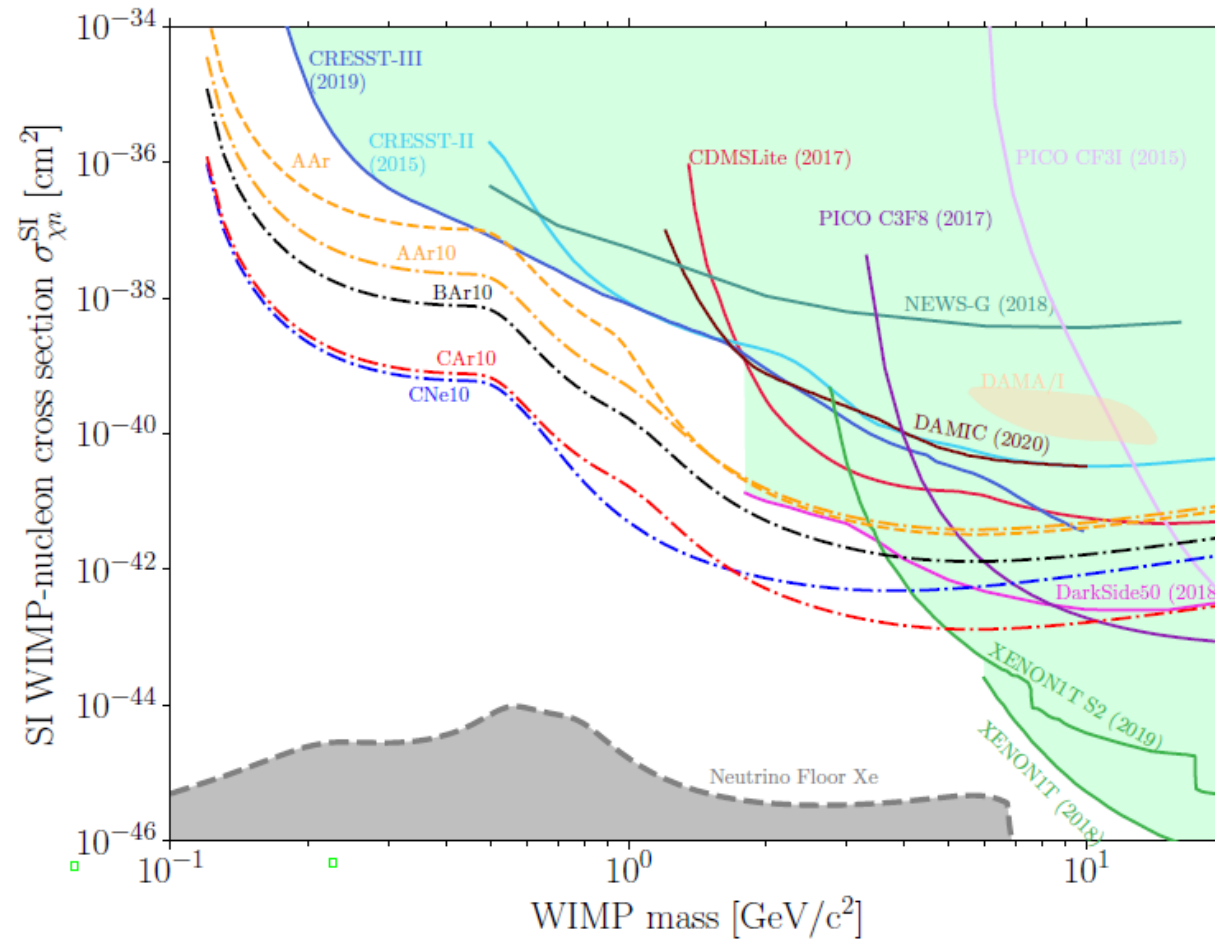


**Fig. 4** Gain curves obtained in the characterization of microbulk Micromegas with Ar (top) and Ne (bottom) mixtures using a  $^{109}\text{Cd}$  source. Curves for different pressures up to 10 bar are shown

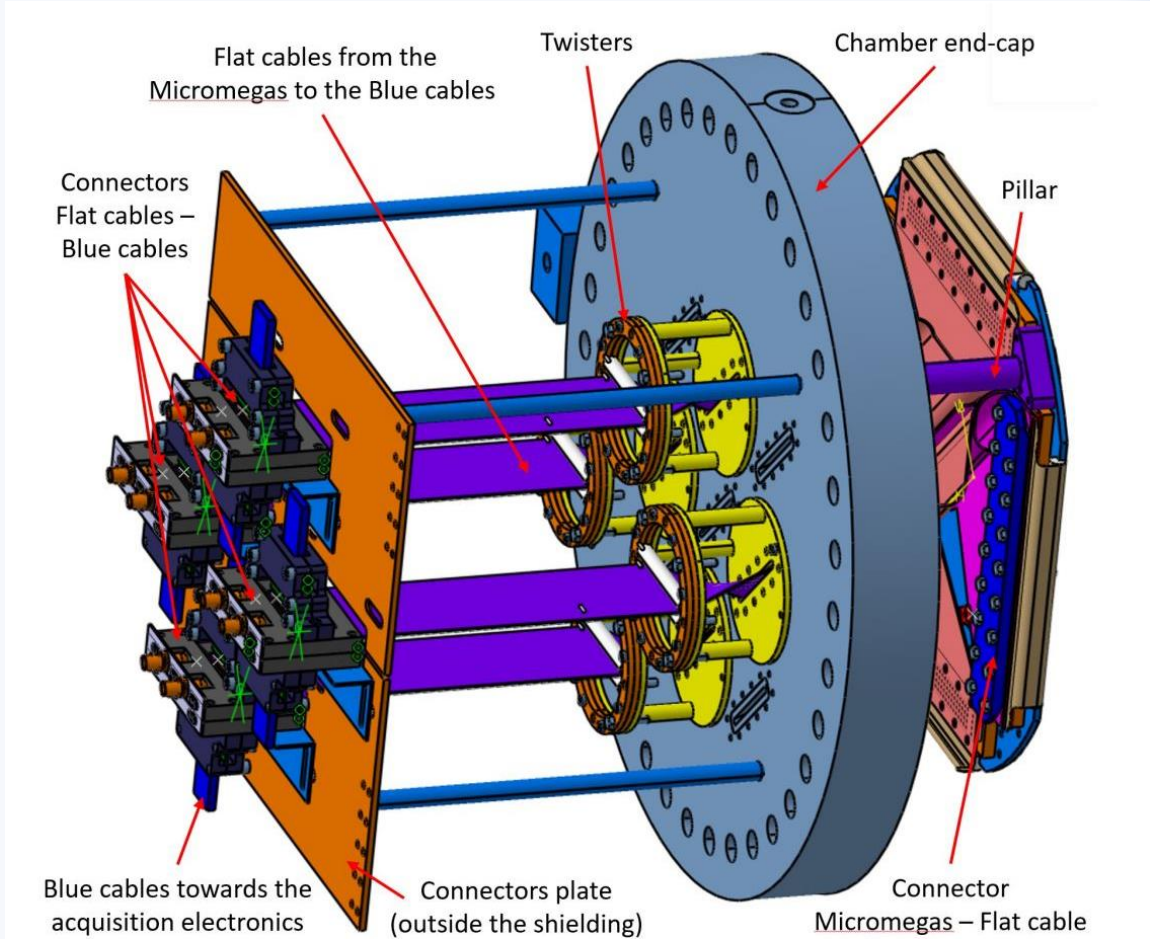
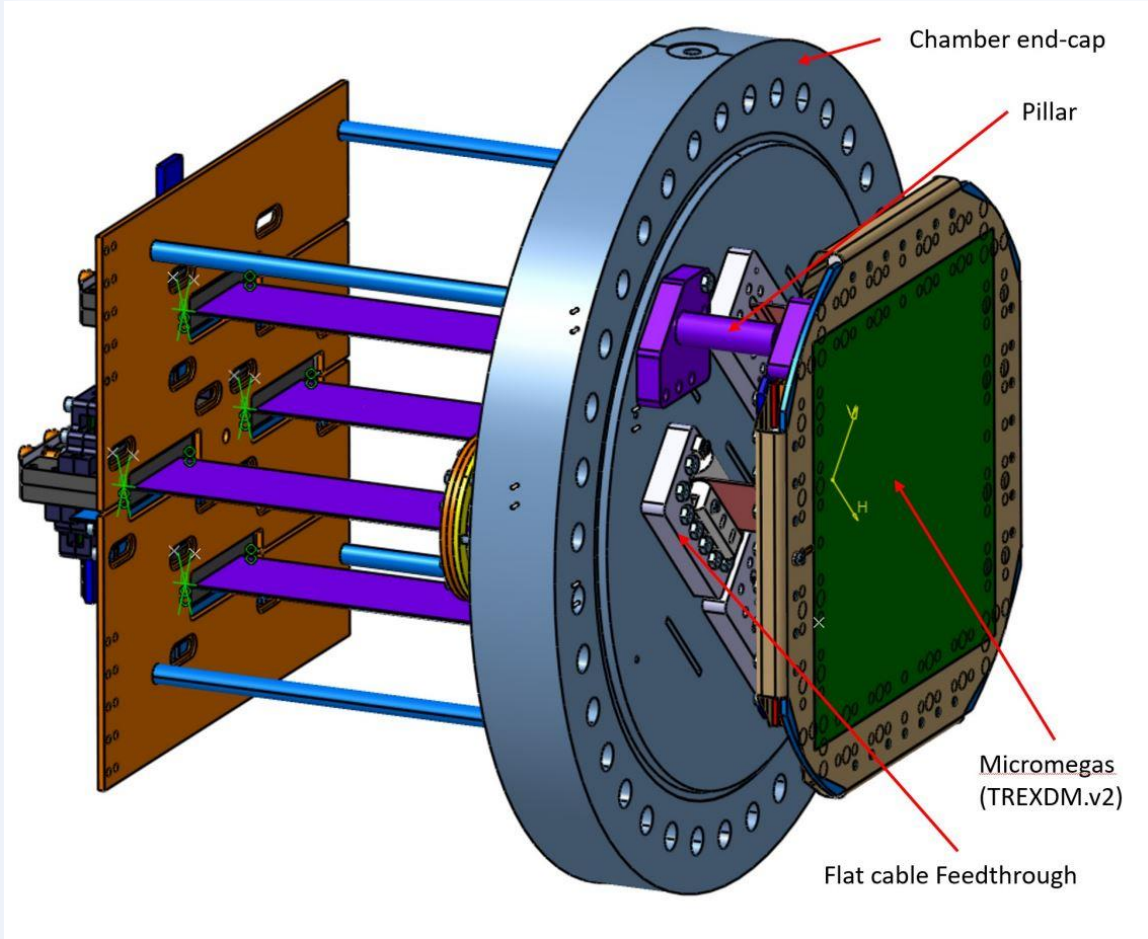
Castel, J., Cebrián, S., Coarasa, I. *et al.* Background assessment for the TREX dark matter experiment. *Eur. Phys. J. C* **79**, 782 (2019).  
<https://doi.org/10.1140/epjc/s10052-019-7282-6>

# Projected sensitivity

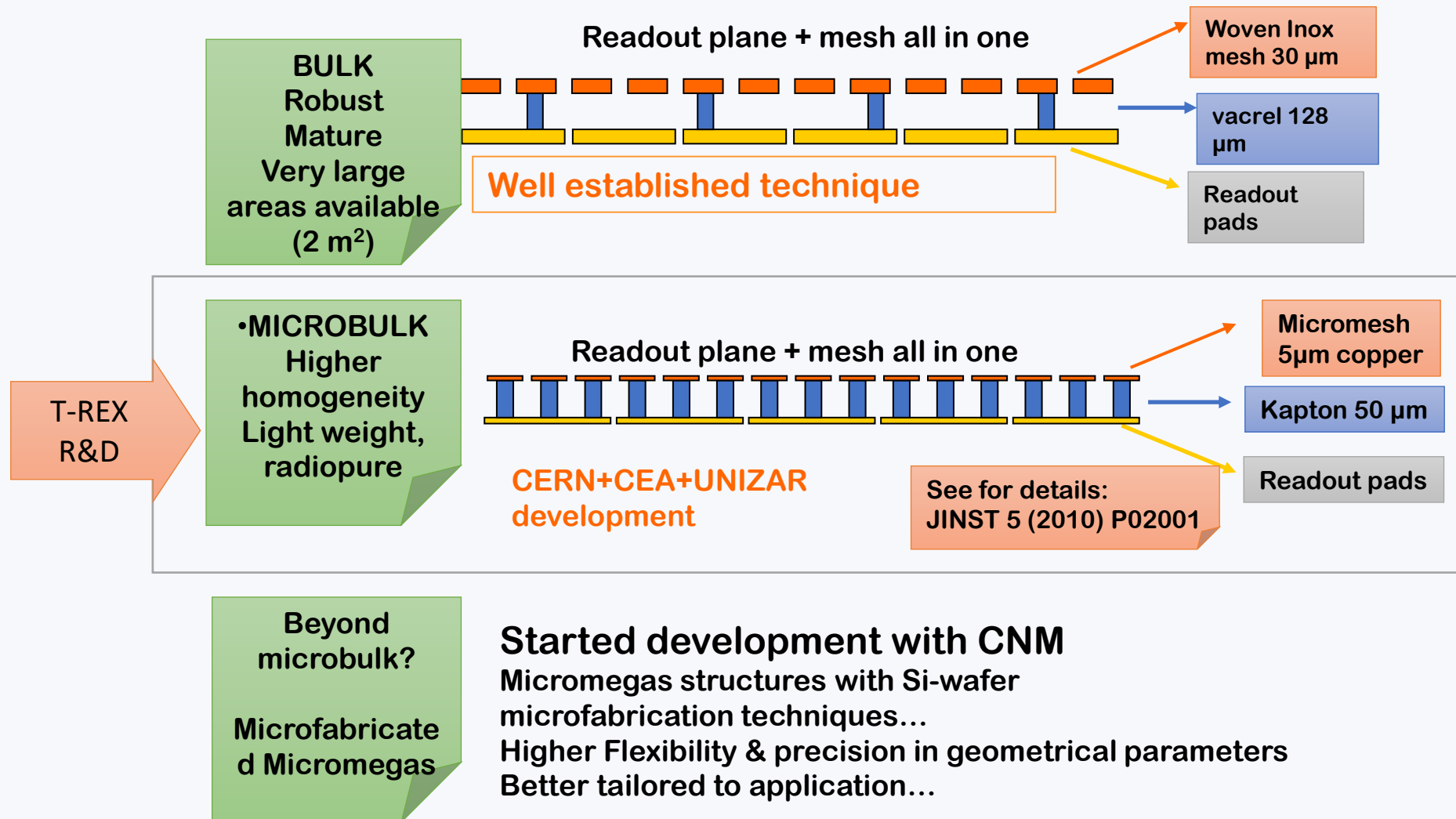
	$E_{th}$ (keV)	$B$ (keV $^{-1}$ kg $^{-1}$ d $^{-1}$ )	Exp (kg y)	Gas
AAr	0.05	10	0.58	Ar-2%Iso
AAr10	0.05	10	0.52	Ar-10%Iso
BAr10	0.05	1	0.52	Ar-10%Iso
CAr10	0.05	0.1	5.2	Ar-10%Iso
CNe10	0.05	0.1	3	Ne-10%Iso



Credit:  
*H. Mirallas*



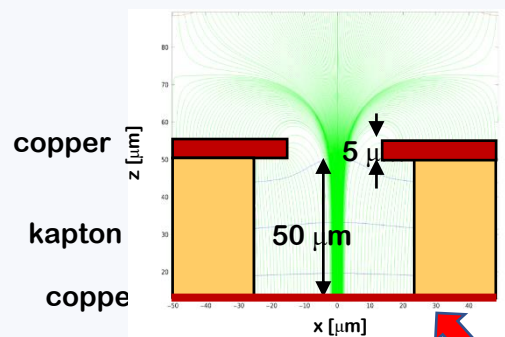
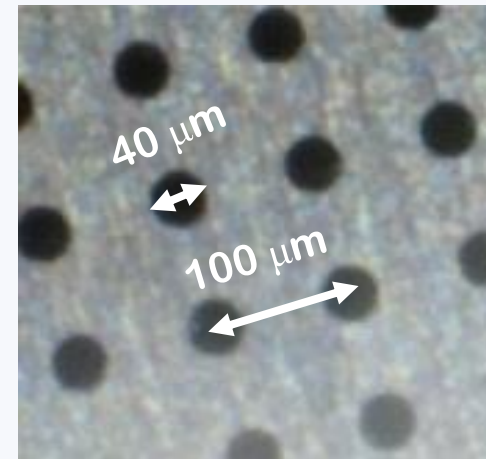
# Micromegas readouts



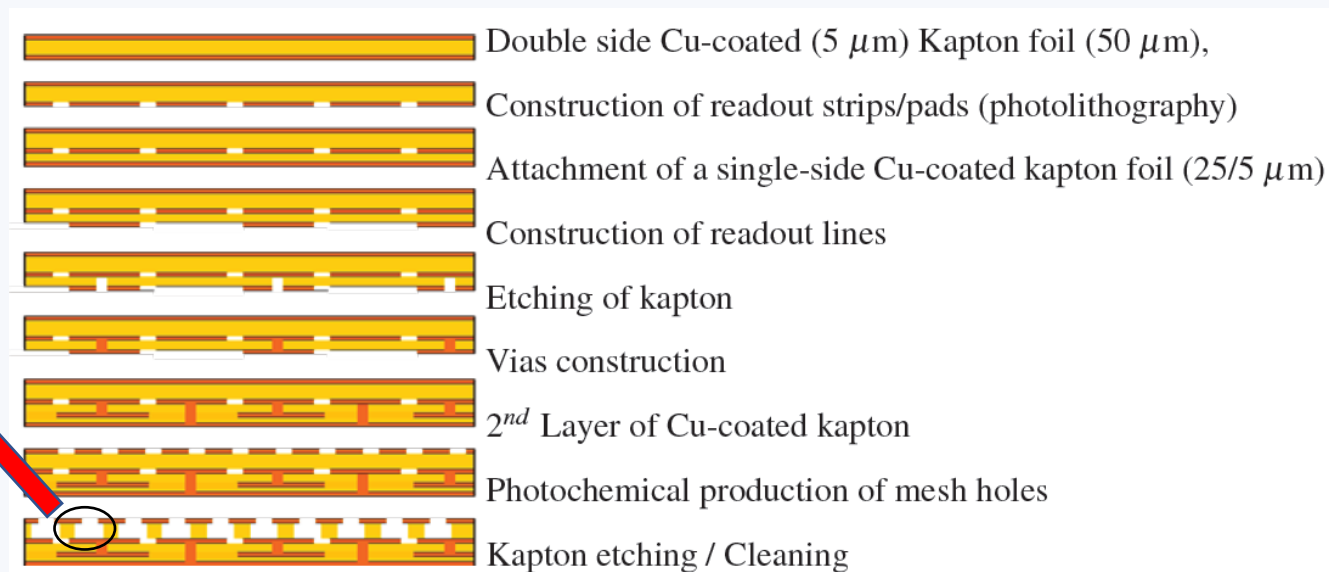
# Microbulk Micromegas

- Made out of copper & polyimide (kapton)
  - potentially very radiopure
- High gap homogeneity
  - good energy resolution
  - Stability/homegeneity in response

Manufactured at Rui de Oliveira's workshop at CERN



See for details:  
JINST 5 (2010) P02001



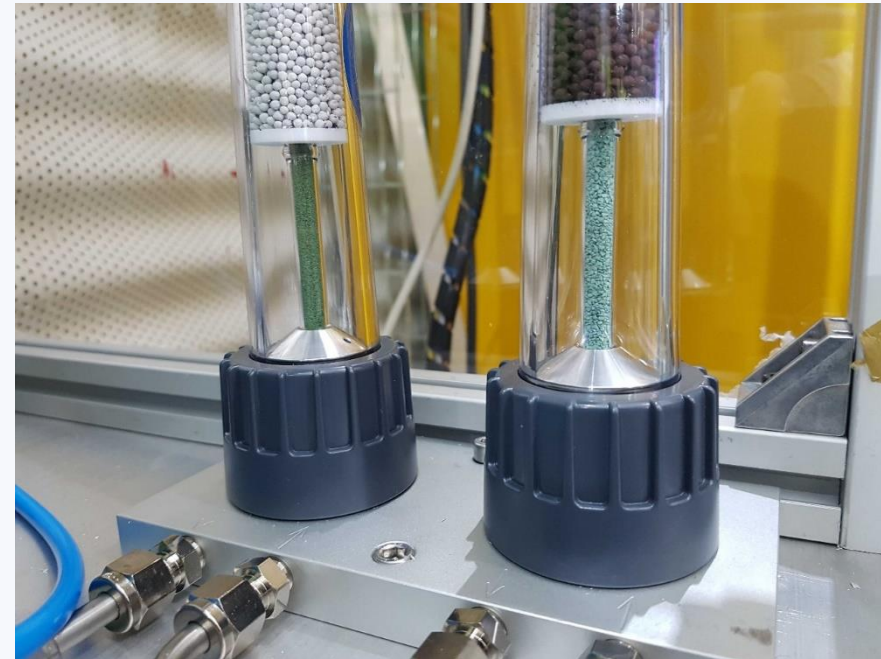
Credit:  
*O.Pérez*

# Background status: Rn issue

Filters in the gas system

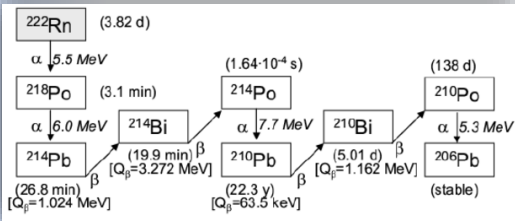


Left: H<sub>2</sub>O; right: O<sub>2</sub>



# Background status: Rn issue

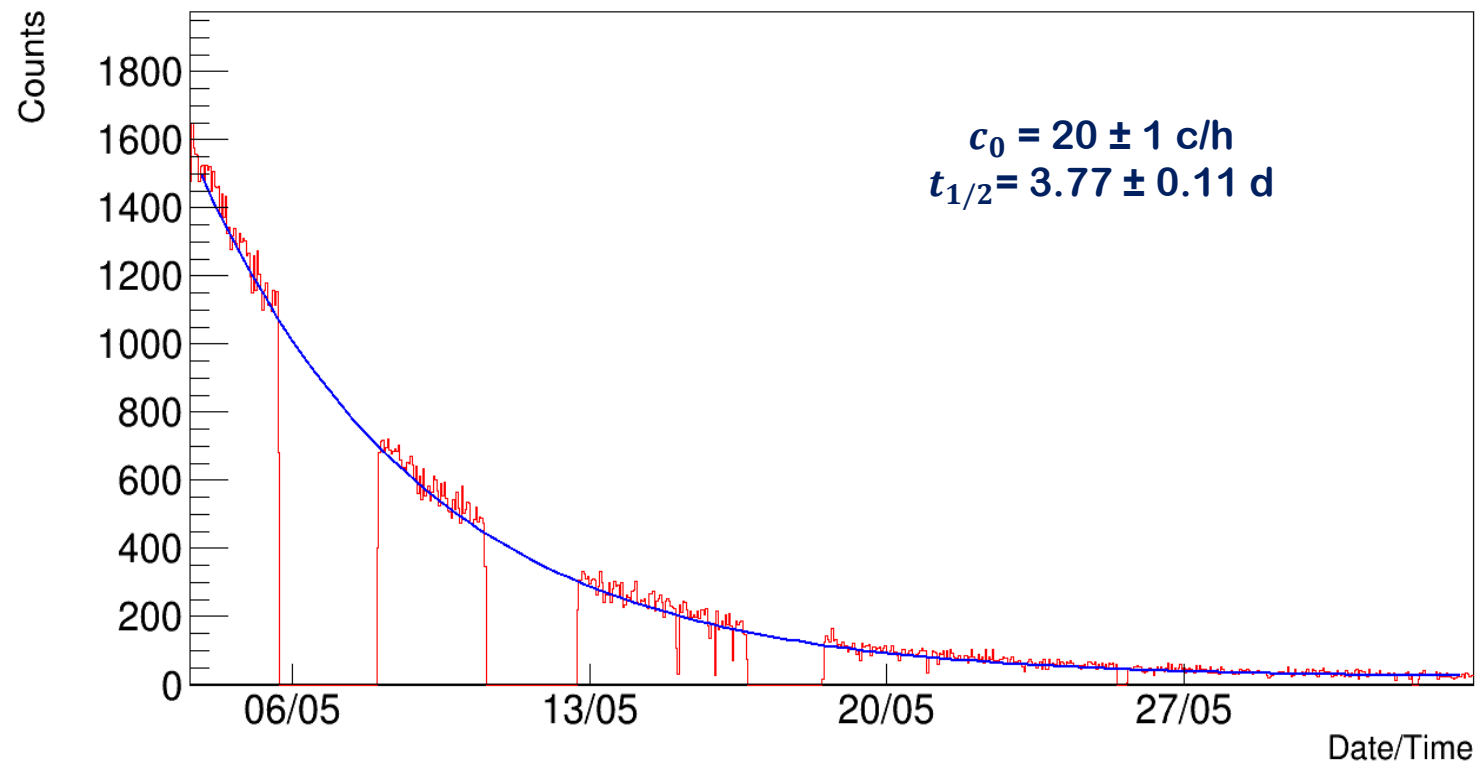
- LE and HE runs during 1 month in seal mode
- They supported the hypothesis of Rn + constant component: decrease in alpha rate and background rate in 0-50 keV after cuts



Low-gain runs (alphas)

$$c_0 + c_1 e^{-c_3 t}$$

Timestamp\_HECuts

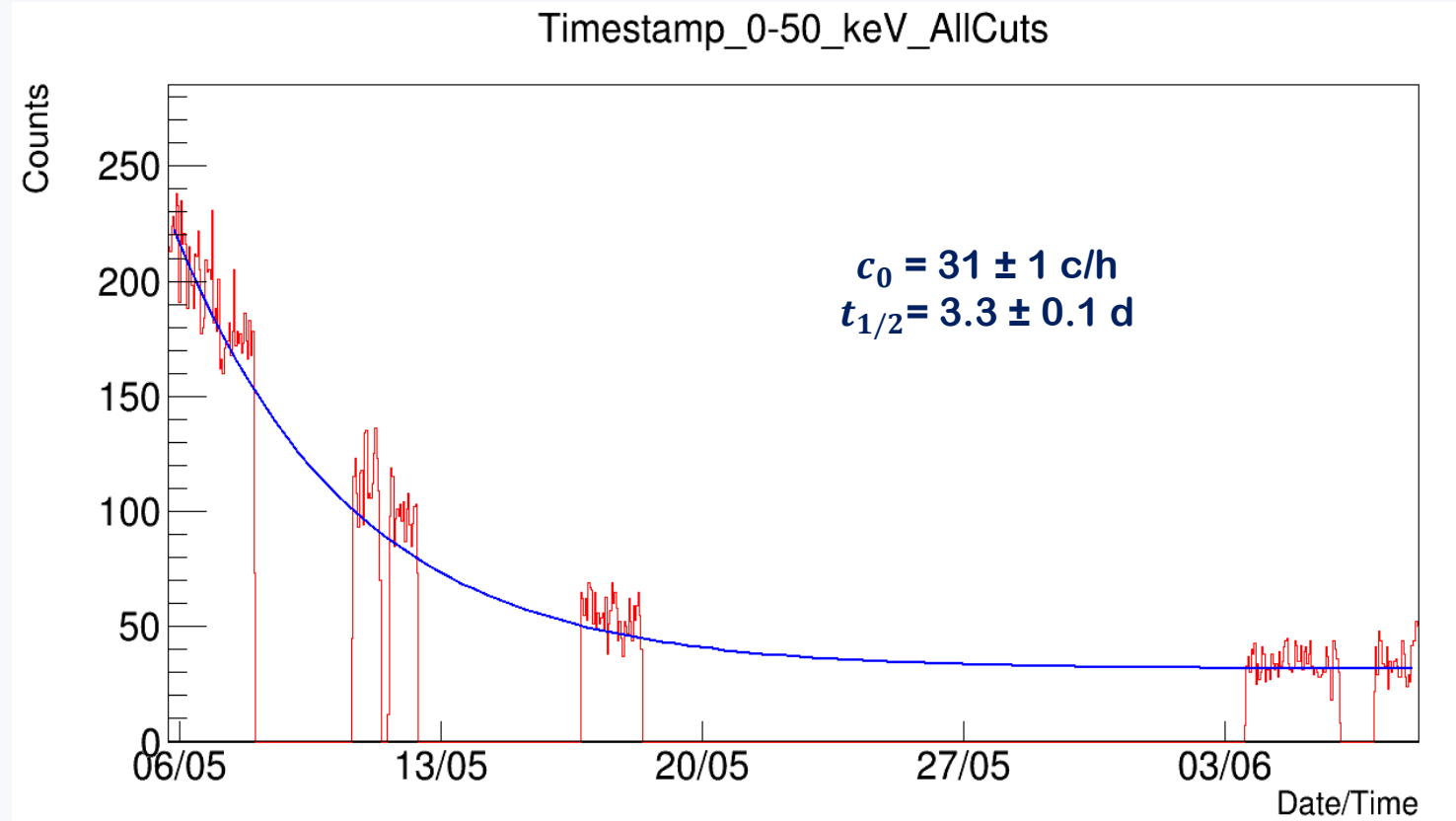




# Background status: Rn issue

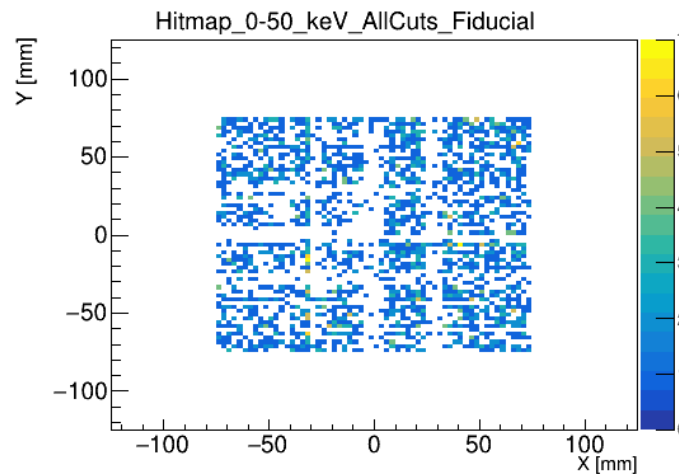
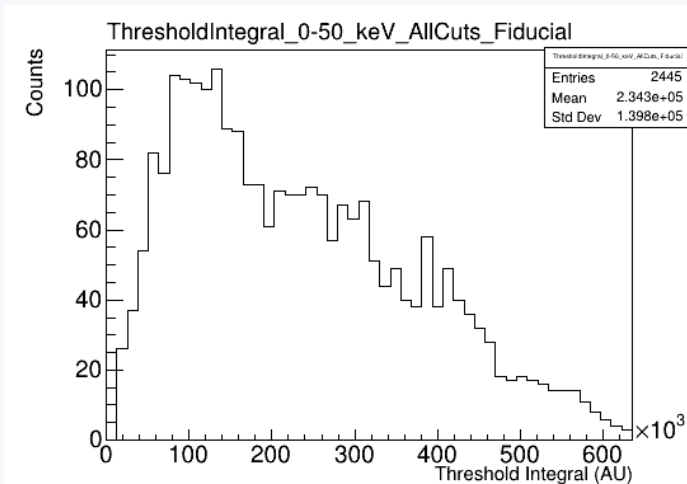
Nominal-gain runs (low energy)

$$c_0 + c_1 e^{-c_3 t}$$

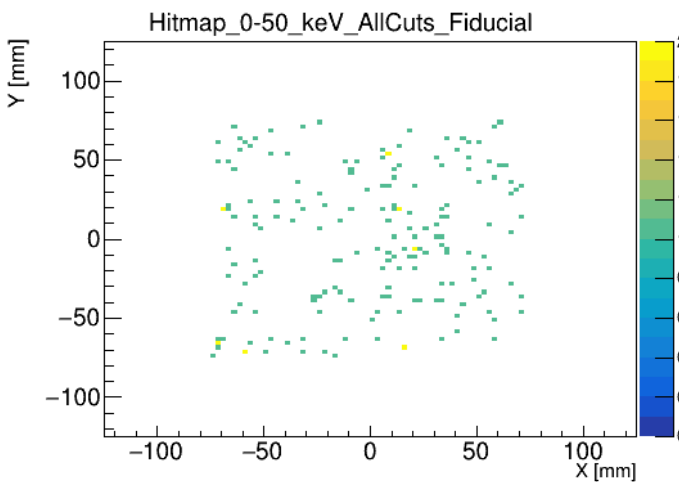
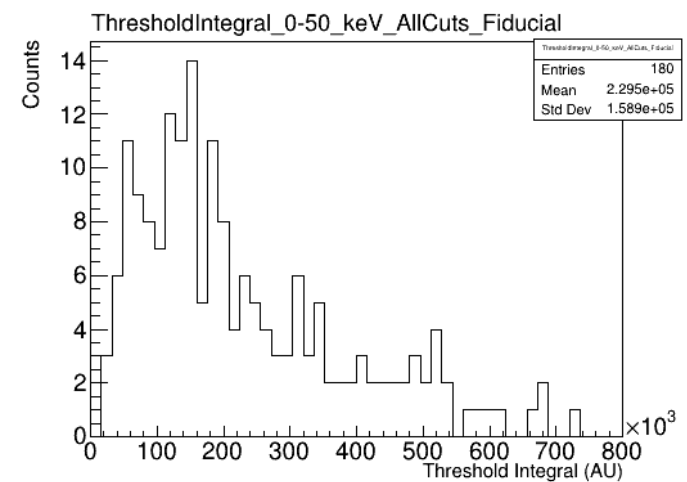


# Background status: Rn issue

## Background @ low energies



Before Rn decay  
0-50 keV  $\rightarrow$  600 dru



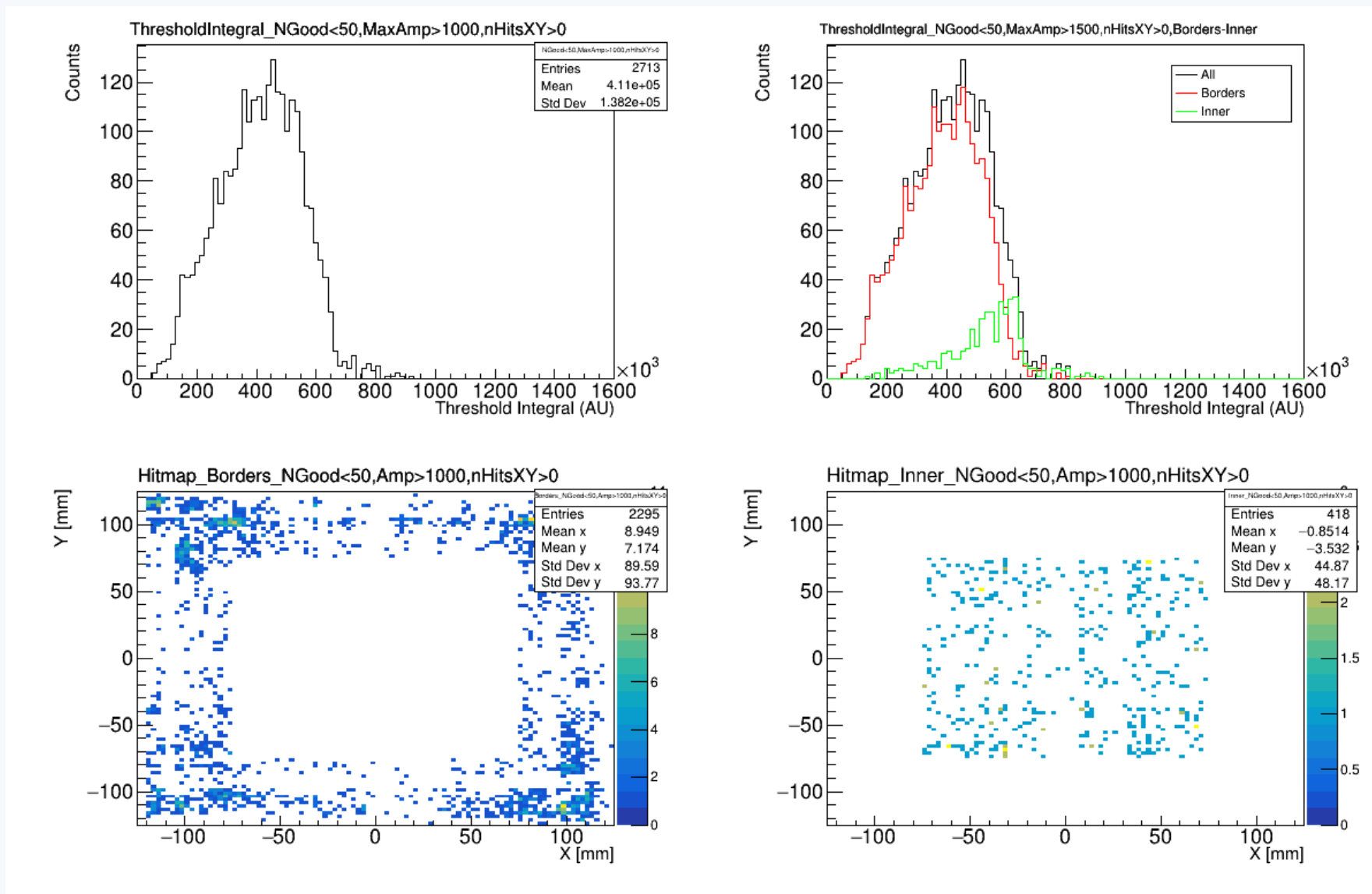
After Rn decay  
0-50 keV  $\rightarrow$  100 dru

# Background status: Rn issue

- Internally emanated Radon is the main source of background (removing it takes us from ~600 dpu down to ~100 dpu in the 0-50 keV range)
- A lot of effort put into removing it from the system:
  - Trying with several commercial filters
  - Testing 5Å molecular sieves (we found out they do trap Rn, but emanate more than Agilent filters, best commercial filters we have)
  - Testing a custom-made O<sub>2</sub>+H<sub>2</sub>O filter developed by the University of Birmingham with low-emanation materials (ongoing collaboration with NEWS-G)
  - Testing activated carbon filters
  - **Open-loop operation bypassing the filters and the recirculation pump**
- Rn progeny surface contamination may well be responsible for the rest of background not accounted for in our background model
  - A program to identify alpha surface contaminations + its mitigation is ongoing

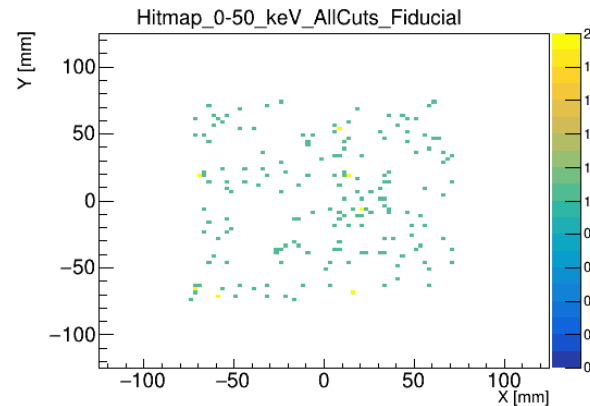
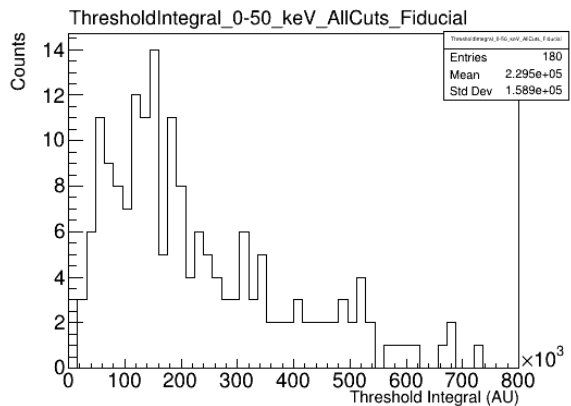
Credit:  
O.Pérez

# Open loop: no longer a Rn spectrum, only surface contamination



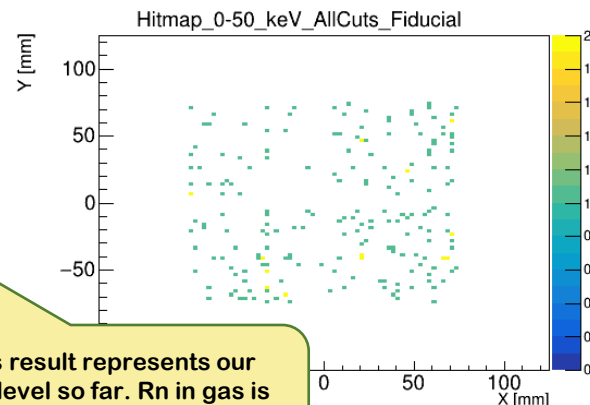
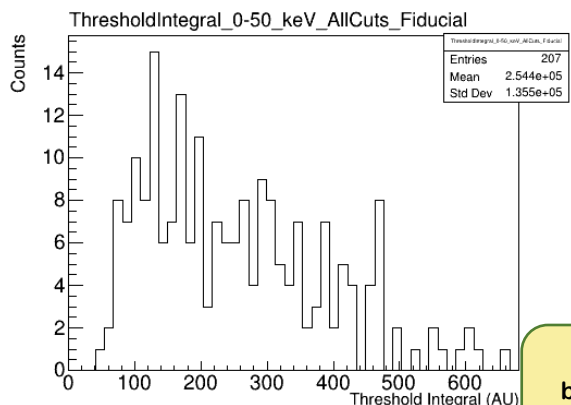
# Background level comparison: seal mode in June 2021 vs. open loop

Sealed mode:  
after Rn decay



0-50 keV: 120 dru

Open loop

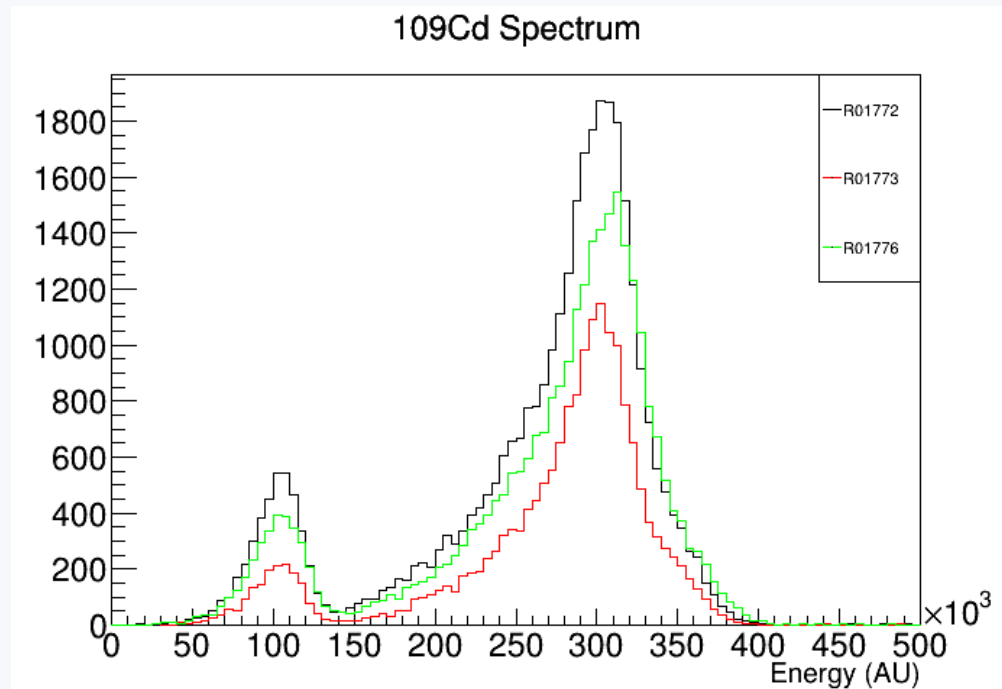


0-50 keV: <80 dru

This result represents our best level so far. Rn in gas is negligible. Now dominated by Pb210 in surfaces

## Flow optimisation in open loop: $^{109}\text{Cd}$ calibrations

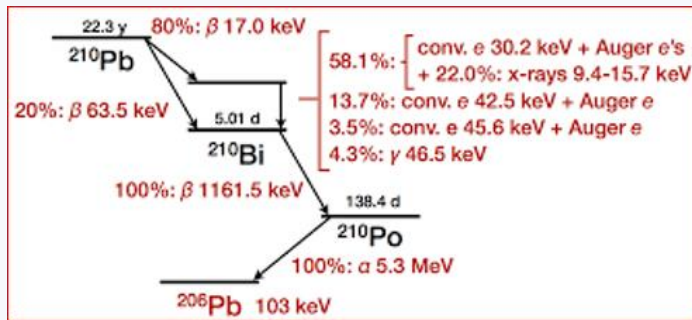
**0.9 ln/h  $\Rightarrow$  one renovation every 11.5 days**



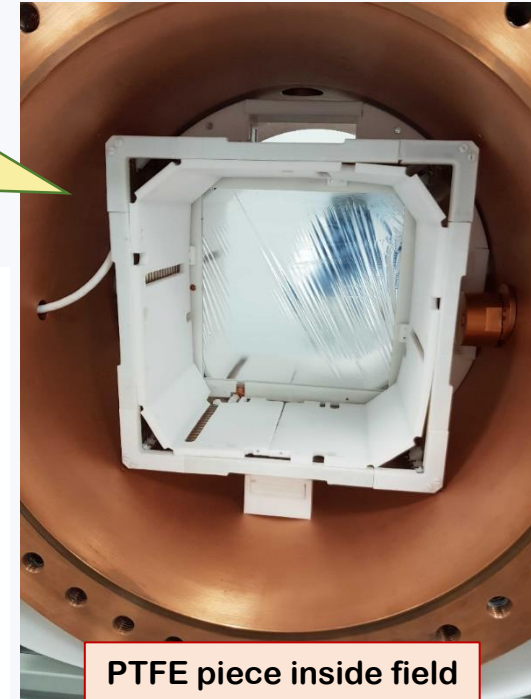
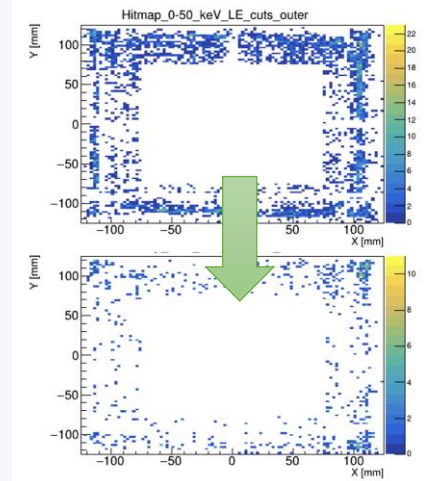
**Calibrations after 0, 1 and 2 weeks at 0.9 ln/h  $\Rightarrow$  gas is clean with this flow**

# Surface alphas (Rn progeny)

- Rn progeny ( $^{210}\text{Pb}$ ) attached to surfaces (from past exposure) produces alpha events, but also LE events (in similar proportion)

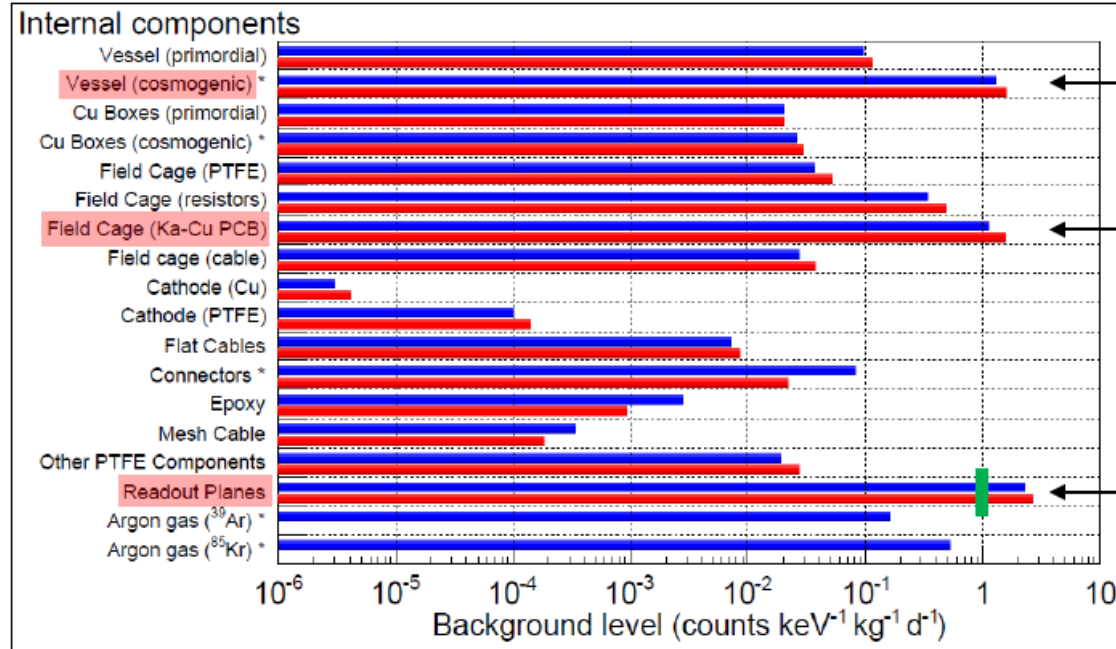


**Reminder:** machining of the Teflon walls led to reduction of both HE and LE background counts in the outer region



PTFE piece inside field cage

# Challenges: Background level reduction



Next step once the background is dominated by the vessel activation.

New design on going.

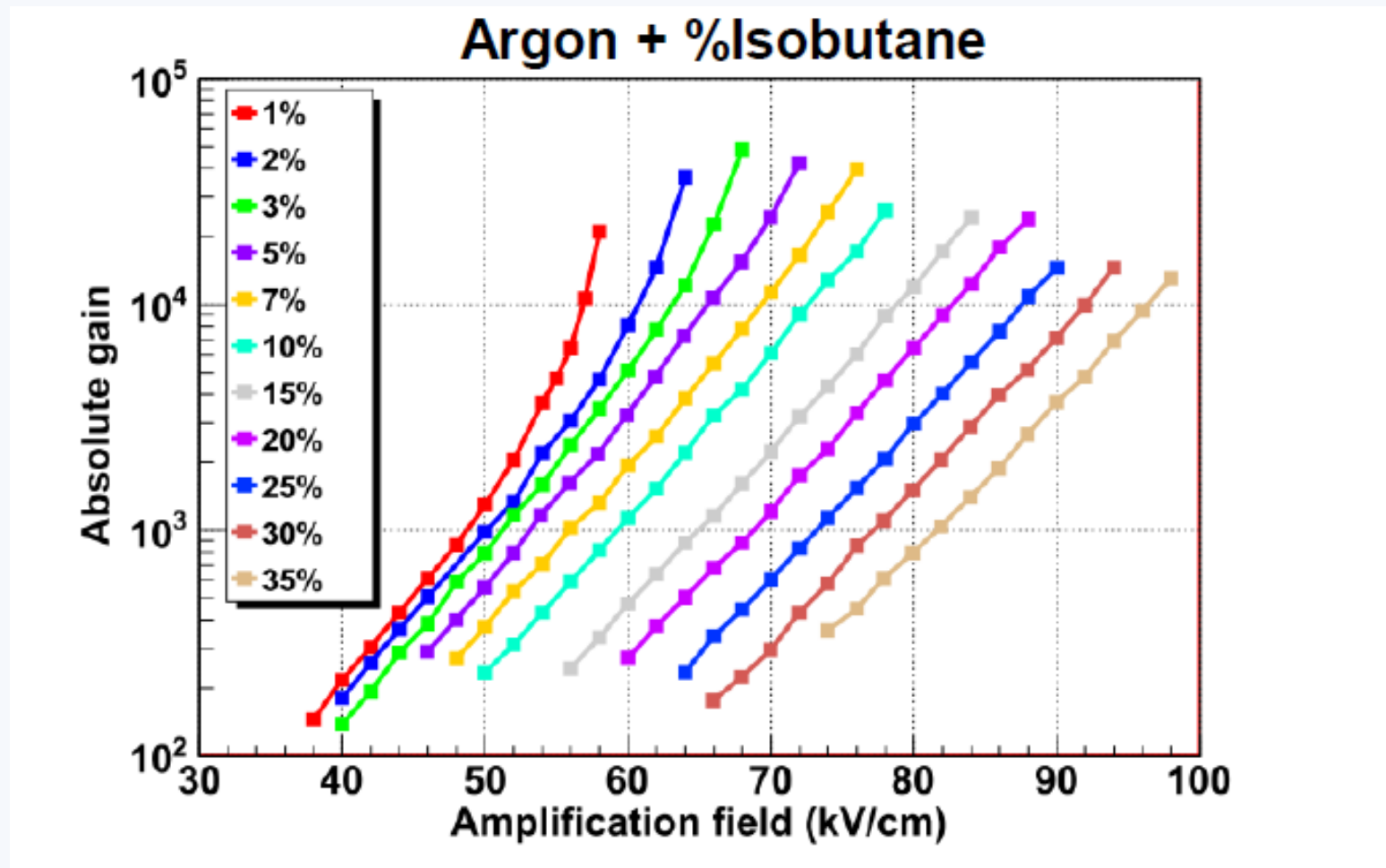
Factor ~ 3 of contamination reduction respect to the previous Micromegas.

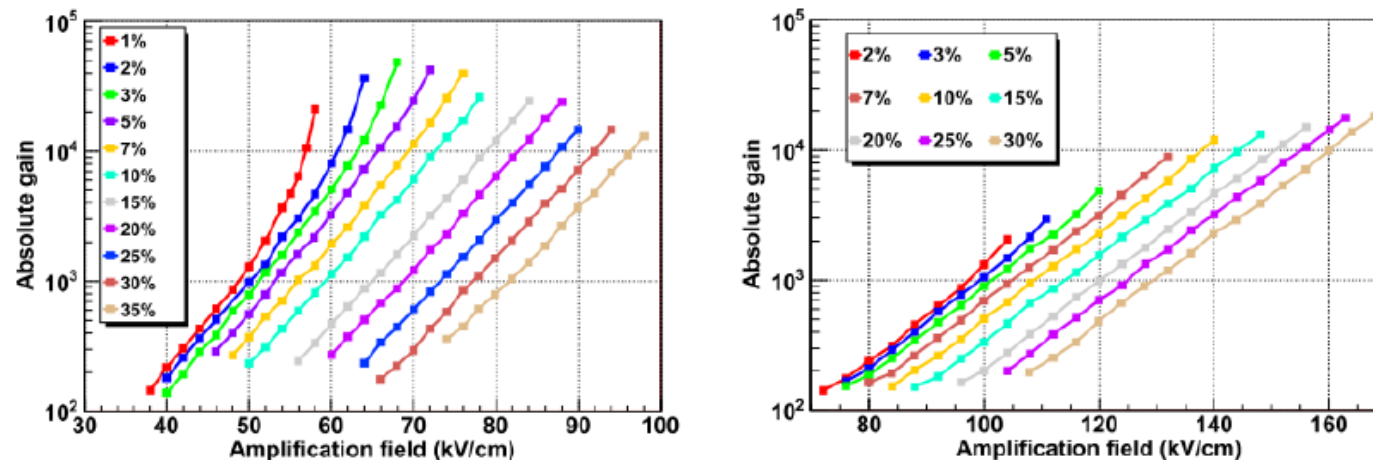
Main challenges in radio-purity of materials:

- Search of clean commercial materials. Large screening programs
- Synthesize clean materials.
- Control of processes in companies
- Storage in controlled environments

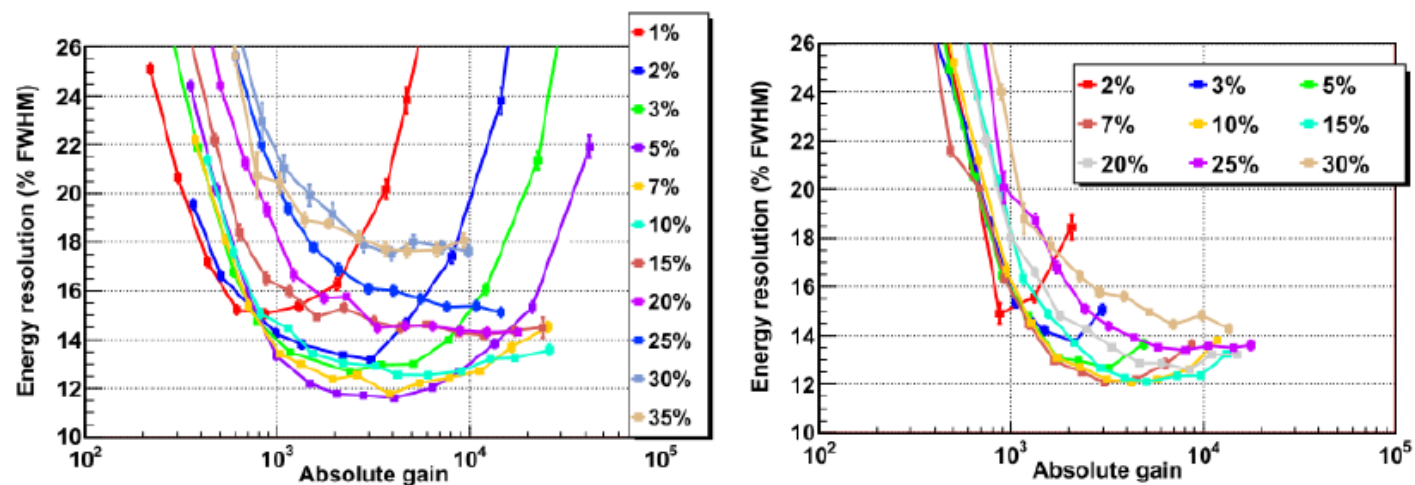


- More isobutane → higher voltages





**Figure 6.** Dependence of the absolute gain with the amplification field for two microbulk detectors with gaps of 50 (left) and 25  $\mu\text{m}$  (right) in argon-isobutane mixtures. The maximum gain of each curve was obtained just before the spark limit. The percentage of each series corresponds to the isobutane concentration.



**Figure 7.** Dependence of the energy resolution with the absolute gain for two detectors of 50 (left) and 25  $\mu\text{m}$ -thickness-gap (right) in argon-isobutane mixtures. The maximum gain of each curve was obtained just before the spark limit. The percentage of each series corresponds to the isobutane concentration.

# Micromegas detector

- Topological information
  - background discrimination

